

MITIGATION OF SELF-GENERATED INTERFERENCE WITH ASKAP PAF

To adopt RFI mitigation techniques, astronomers need to understand the effects of the algorithms on astronomical figures of merit.

CURRENT RFI MITIGATION

- Corrupted data is flagged (discarded) in frequency and time, reducing sensitivity.
- Not using the spatial information in the system.
- Implementing spatial filtering algorithms can reduce the amount of flagged data

UNWANTED SIGNAL (STATIC)

- Due to clock signal in the ASKAP receiver electronics.
- Perform mitigation once at the start of an observation.

DYNAMIC CASE

- Intermediary use case to mitigating more pernicious forms of moving RFI including navigation satellite and aircraft.
- Mitigation performed continuously.

**MITIGATION
ALGORITHM**

UNWANTED
SIGNAL

BEAMFORMED
WEIGHTS

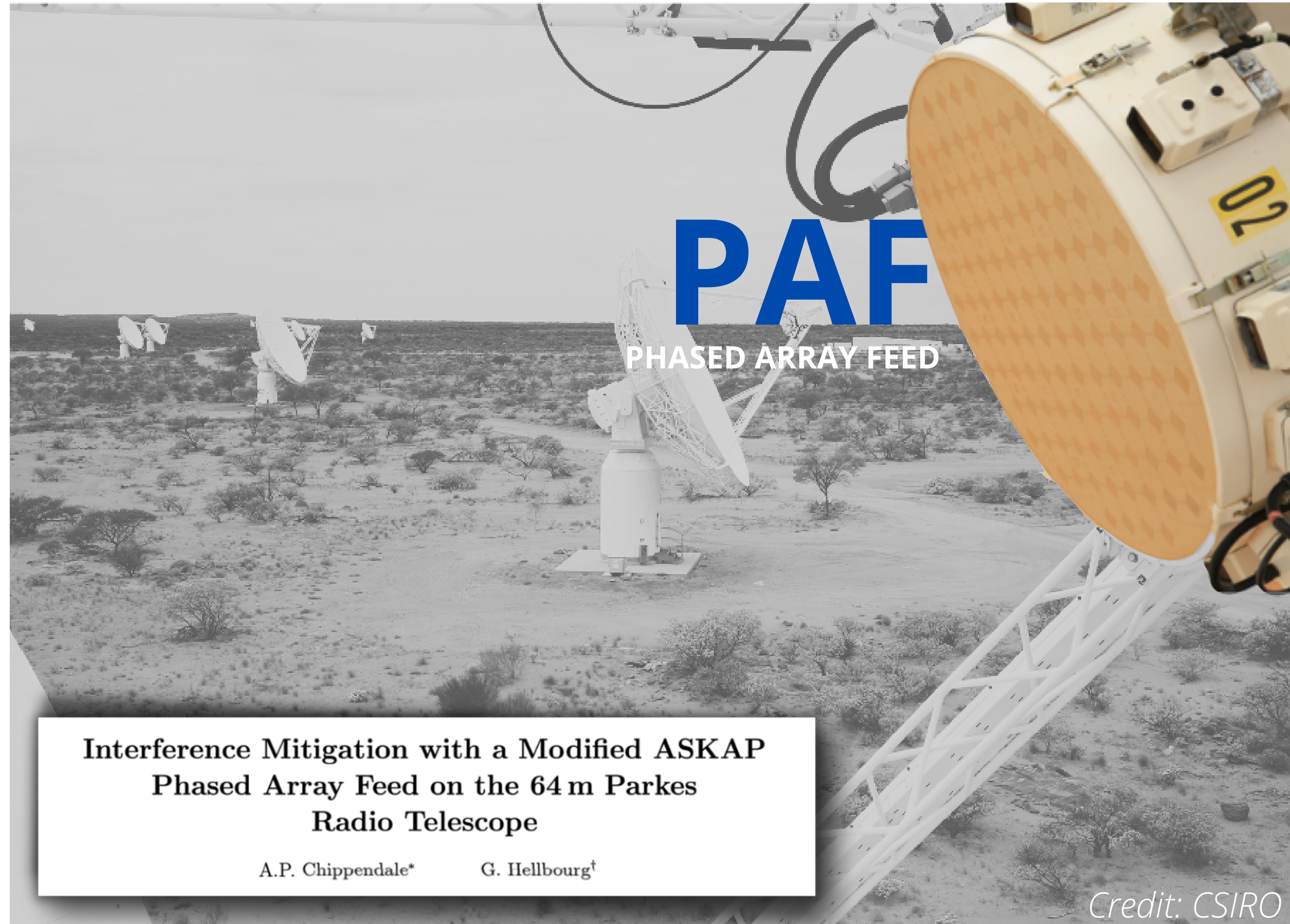
RESULTS

SIMULATED

OFFLINE

ON A SINGLE
ANTENNA

OVERVIEW

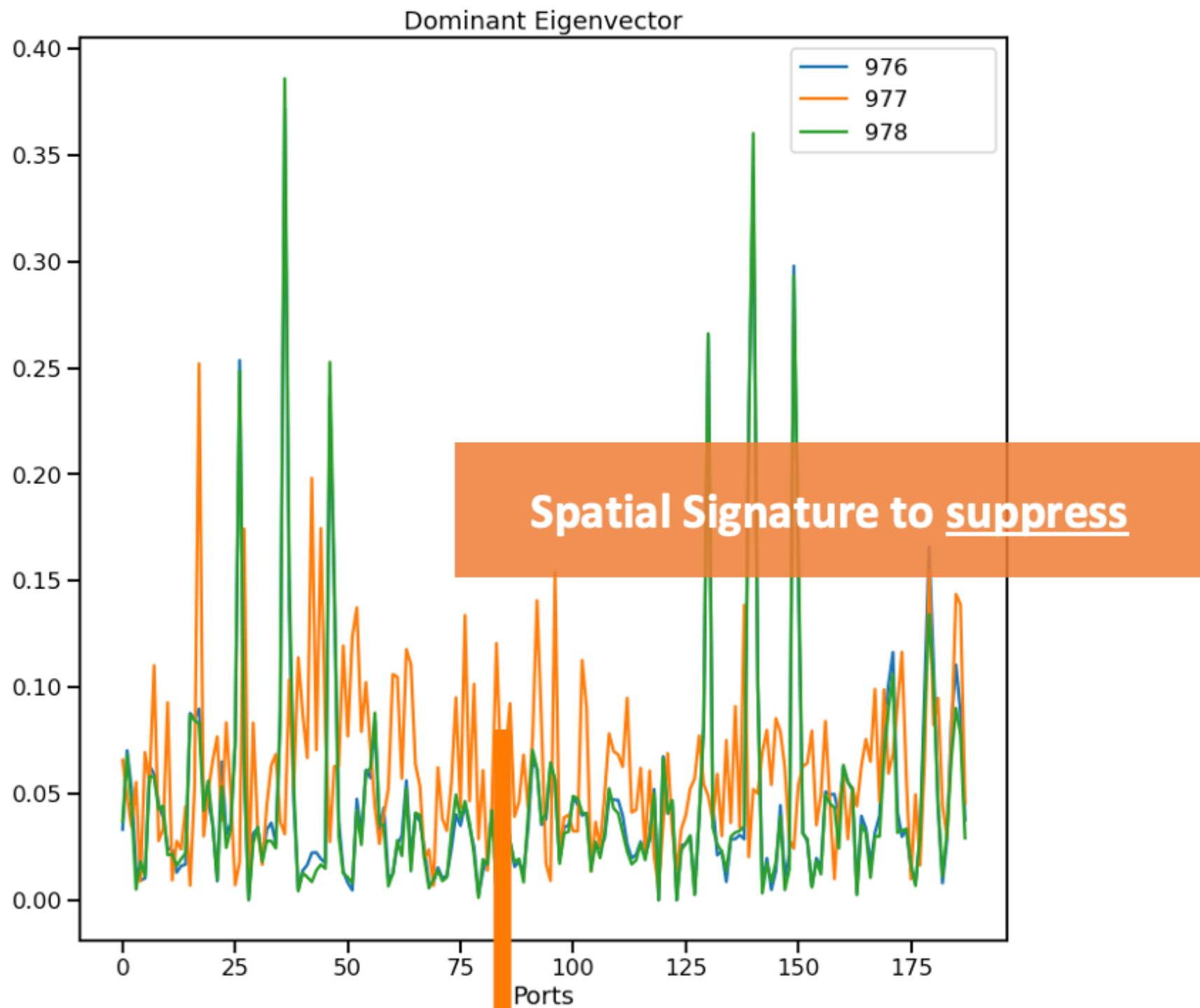


**Interference Mitigation with a Modified ASKAP
Phased Array Feed on the 64 m Parkes
Radio Telescope**

A.P. Chippendale*

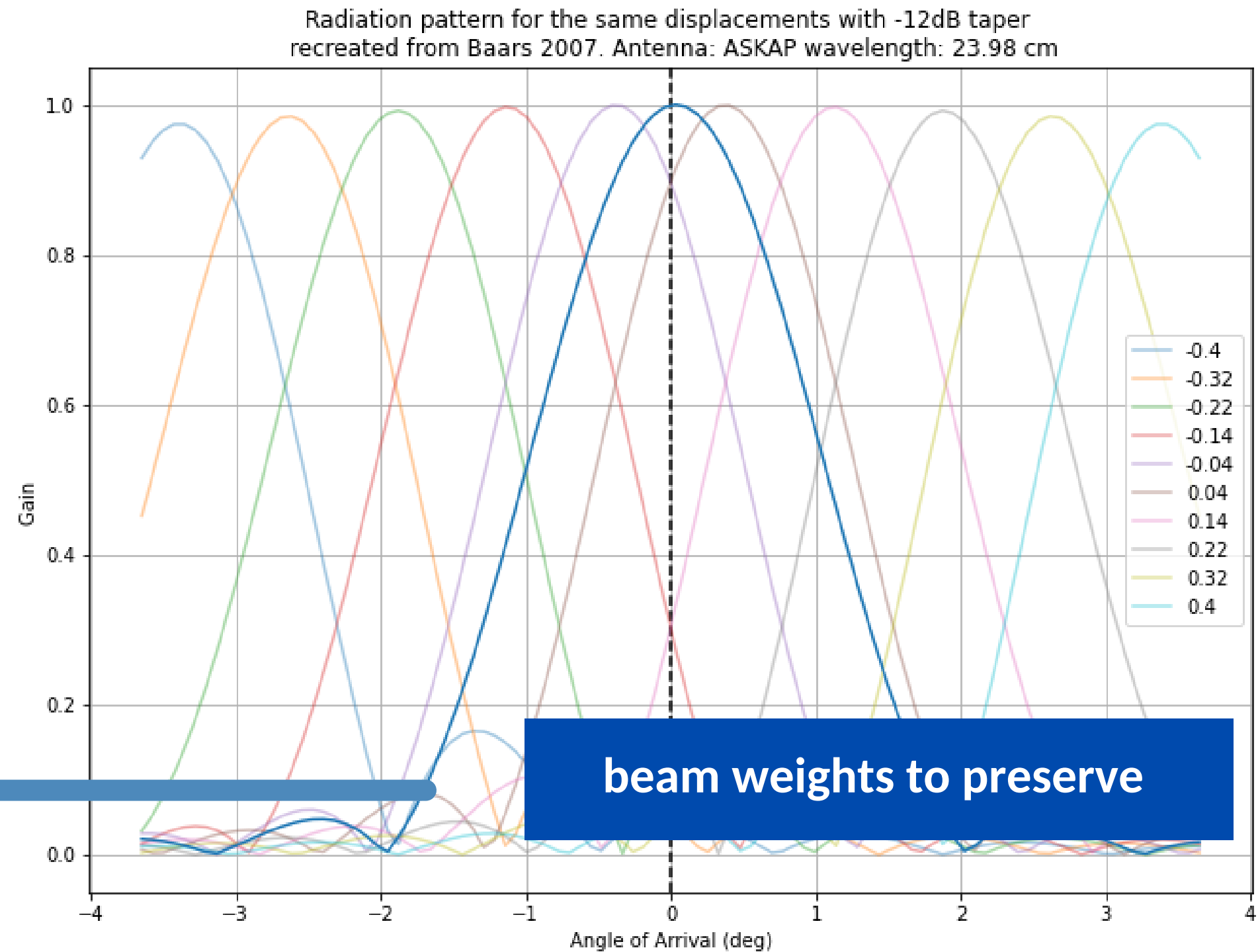
G. Hellbourg†

Credit: CSIRO



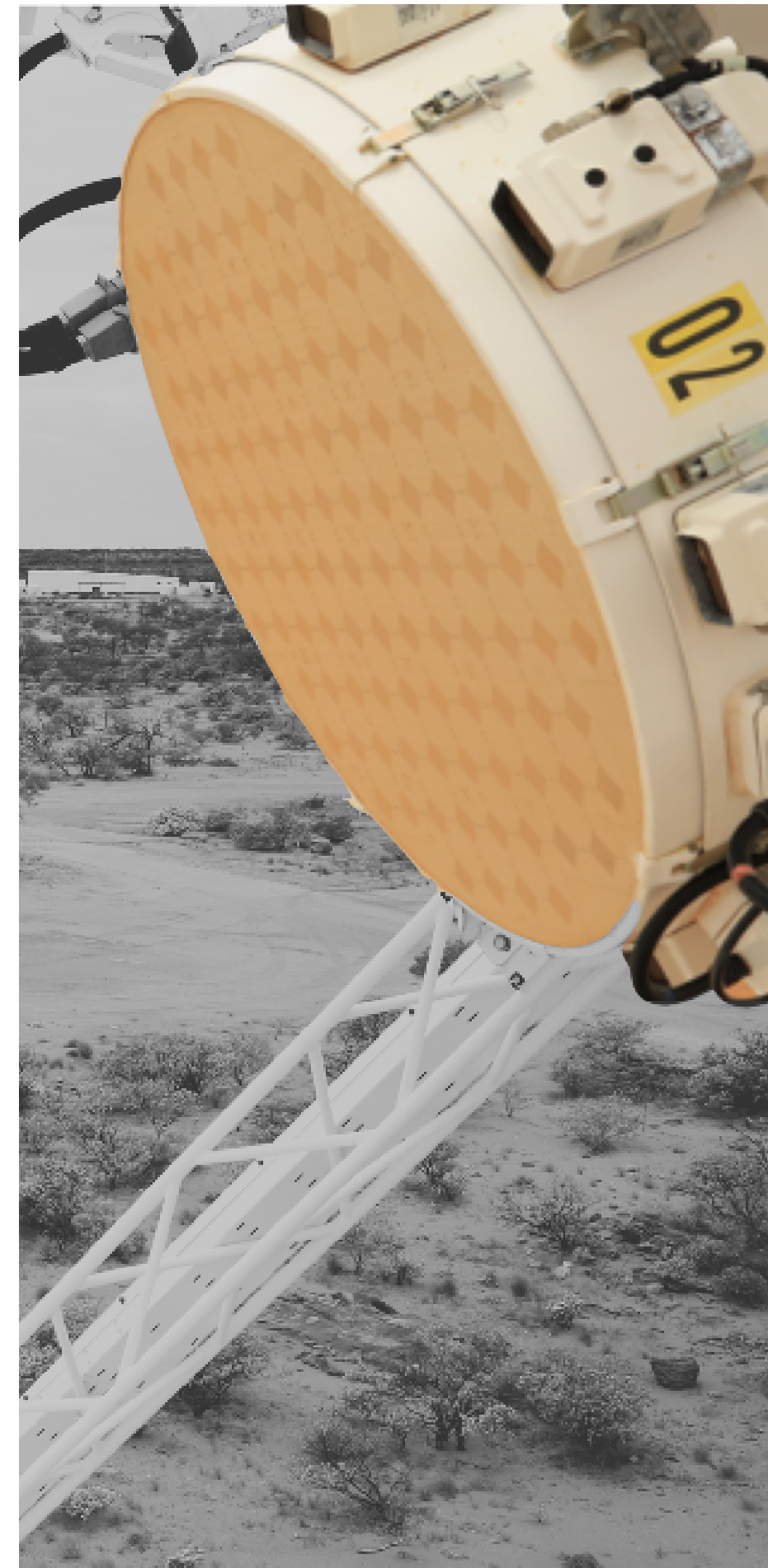
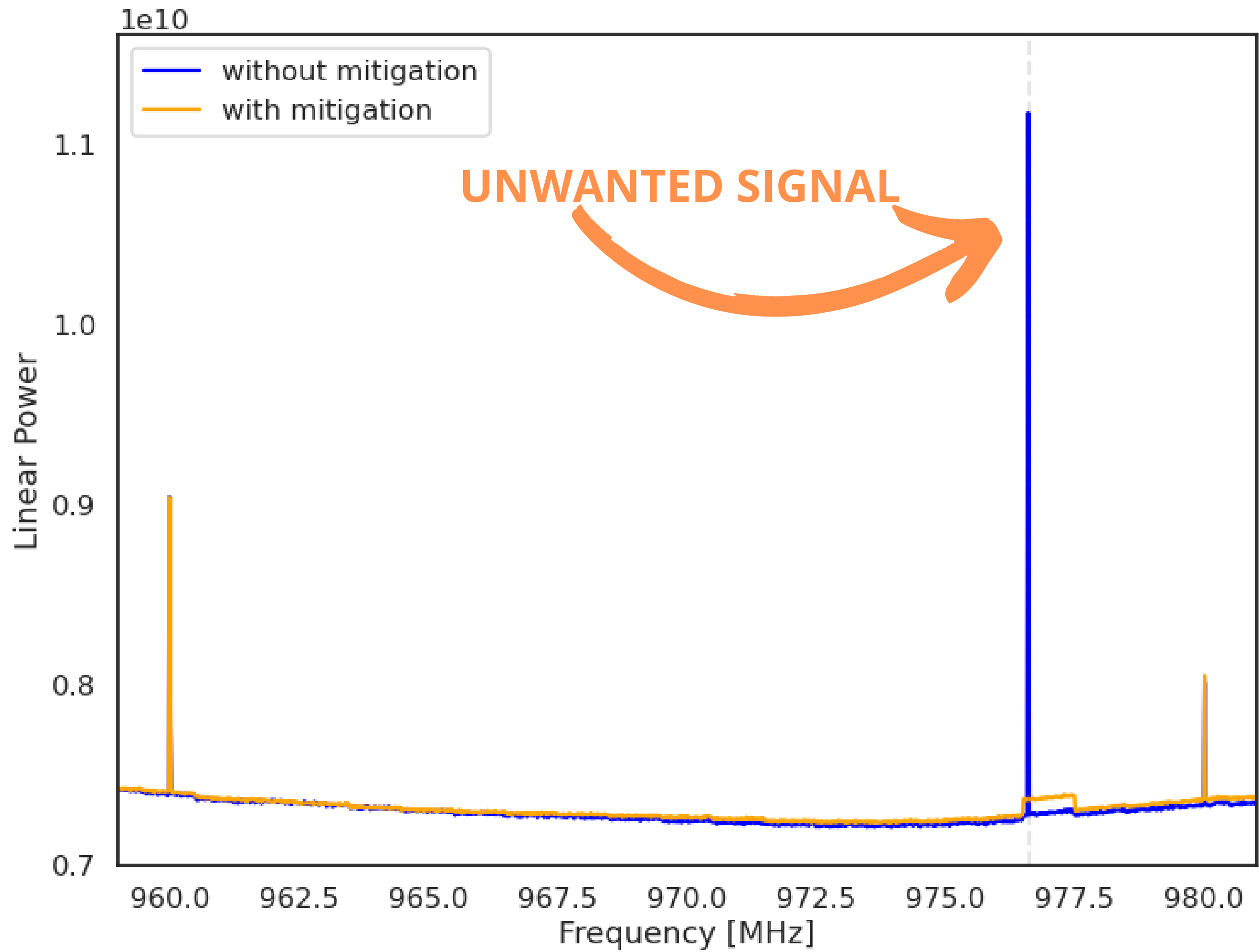
SPATIAL FILTERING EXPERIMENT WITH THE ASKAP BETA ARRAY

Gregory Hellbourg^(1,2), Keith Bannister⁽²⁾, Aidan Hotan⁽²⁾

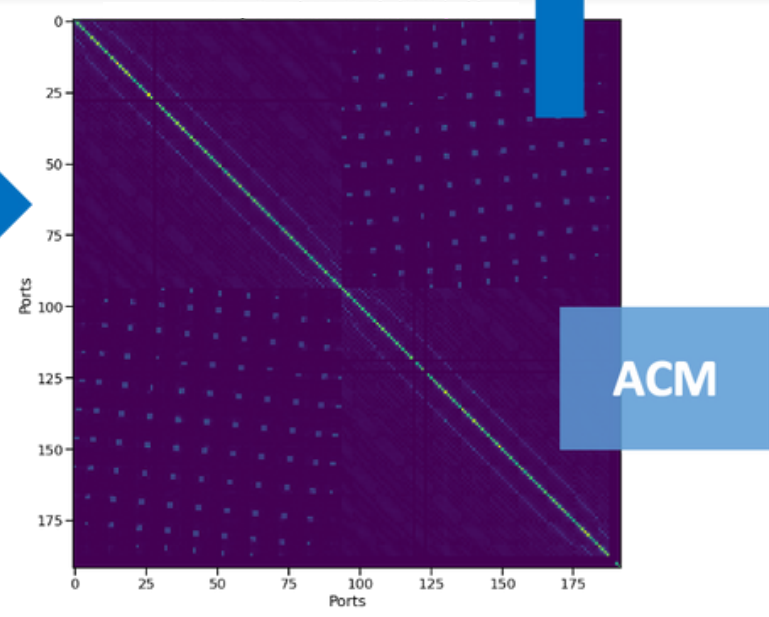
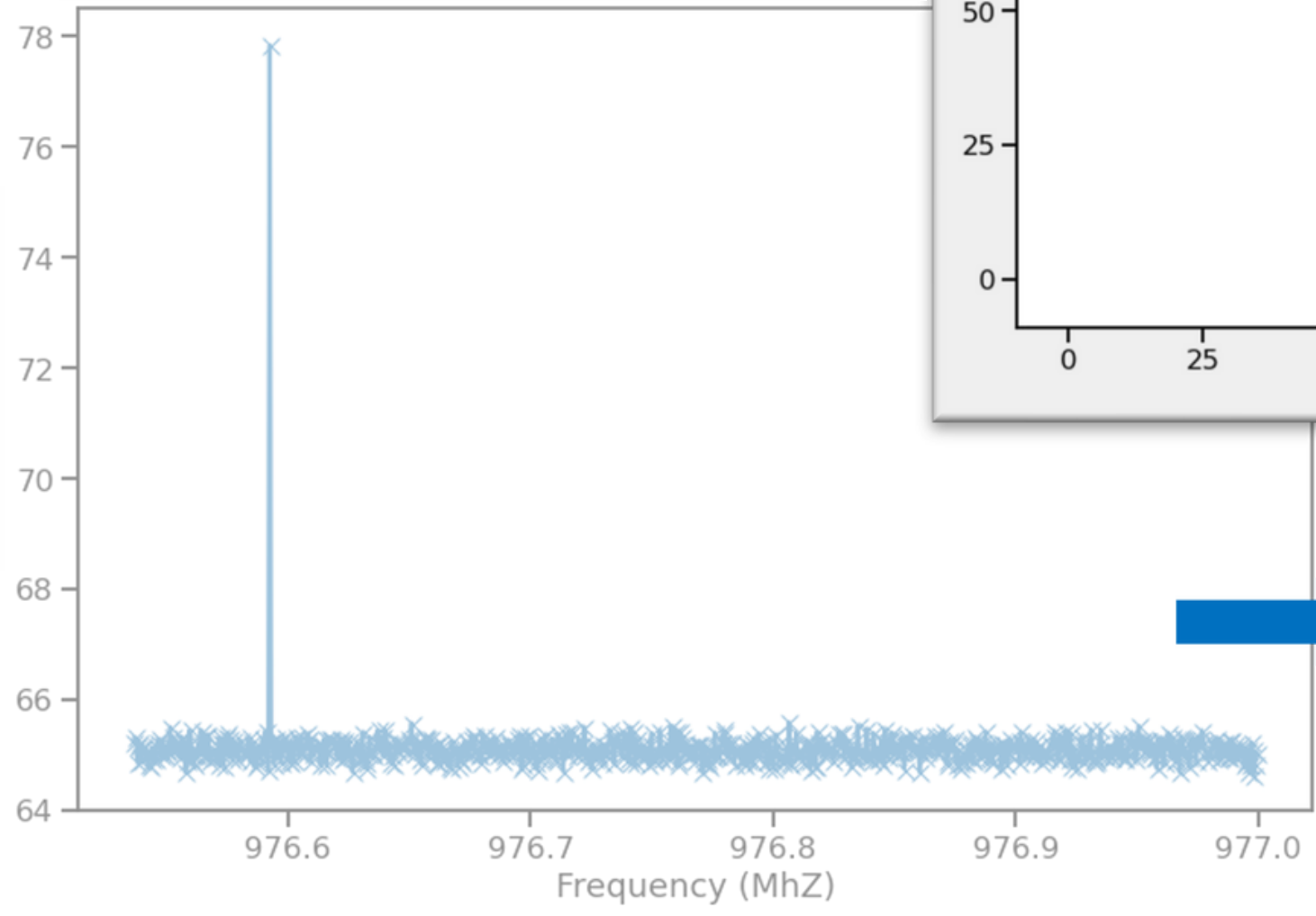
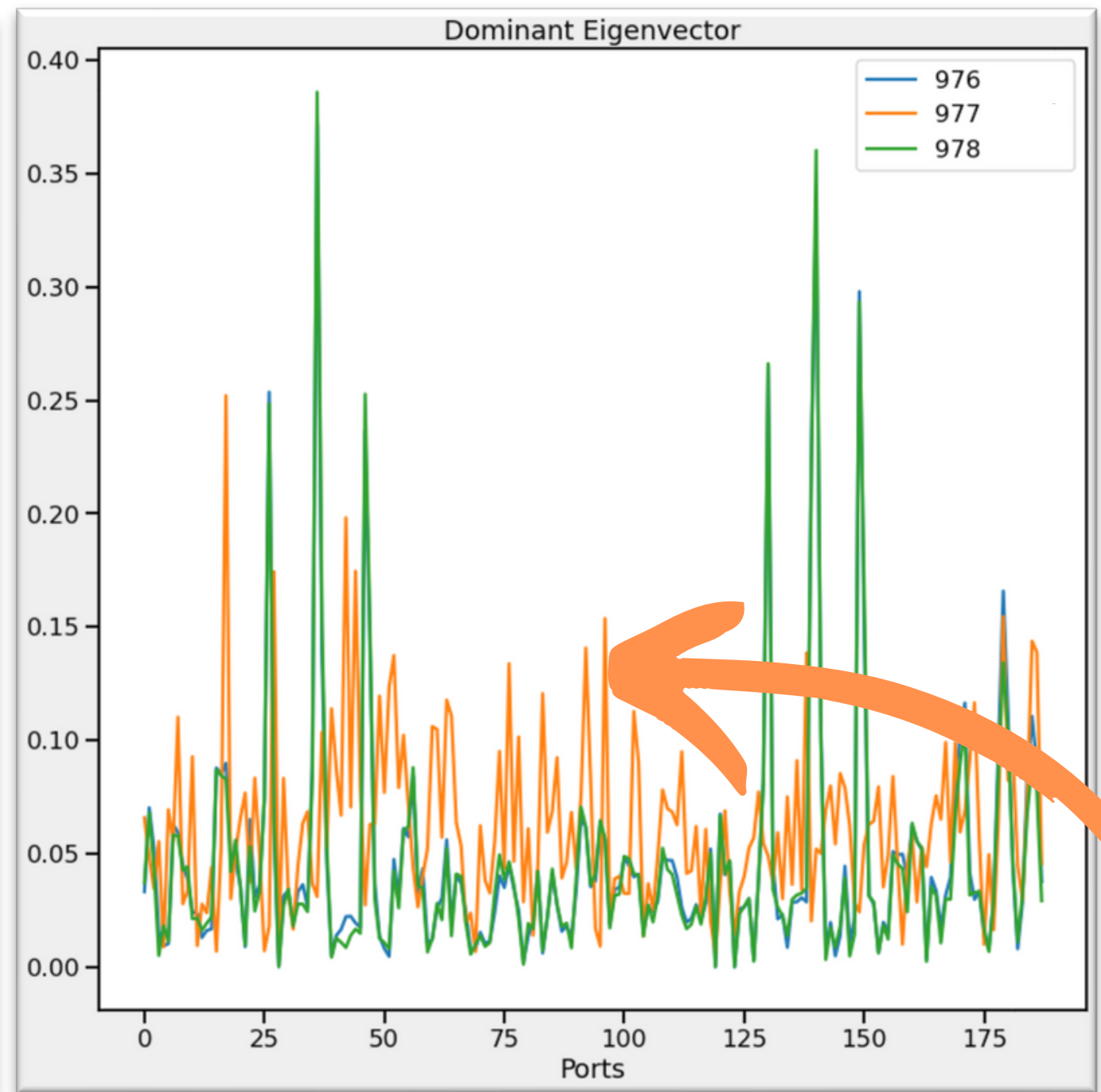
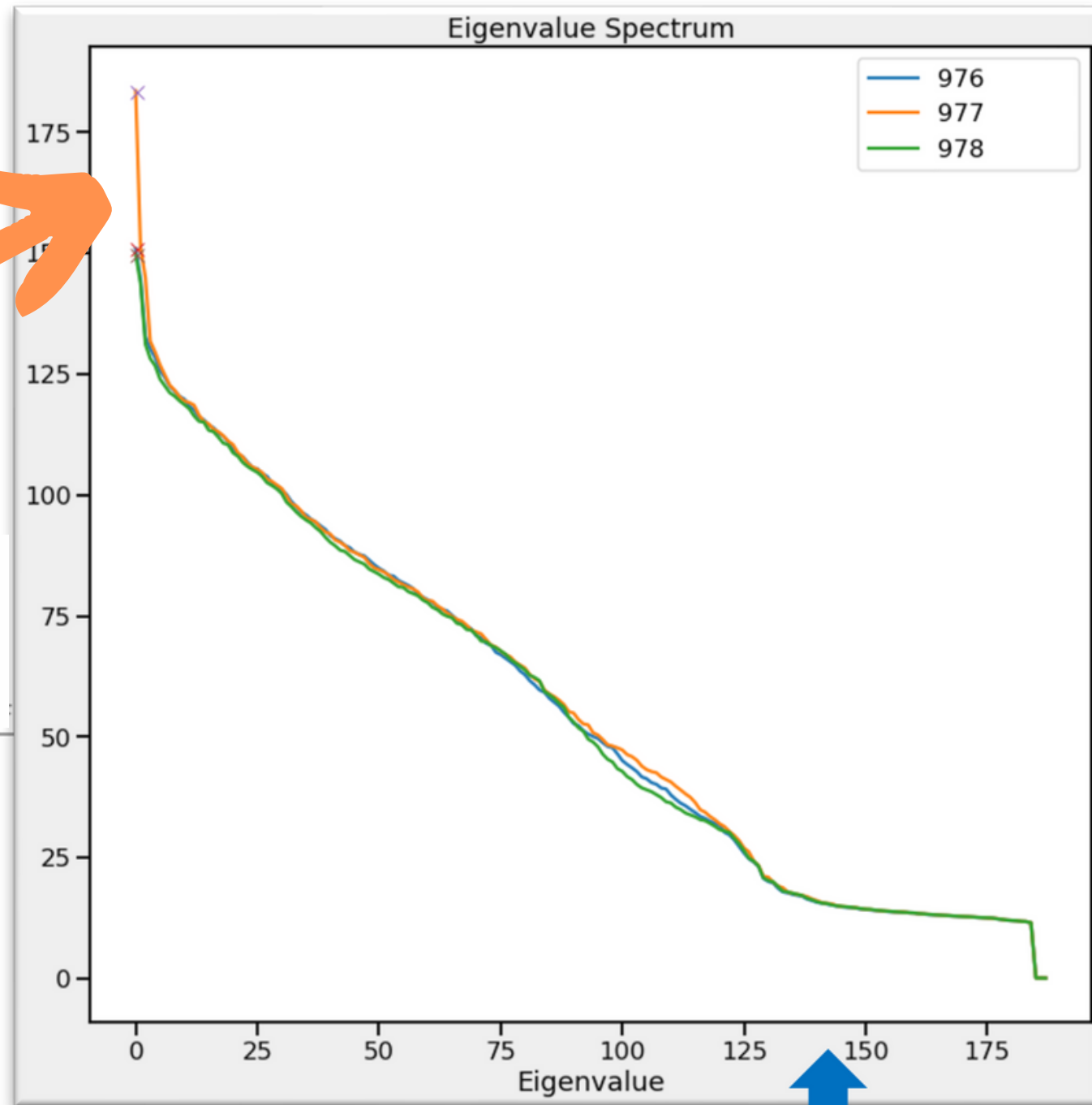


$$P_{ortho} = I - a_r (a_r^H a_r)^{-1} a_r^H$$

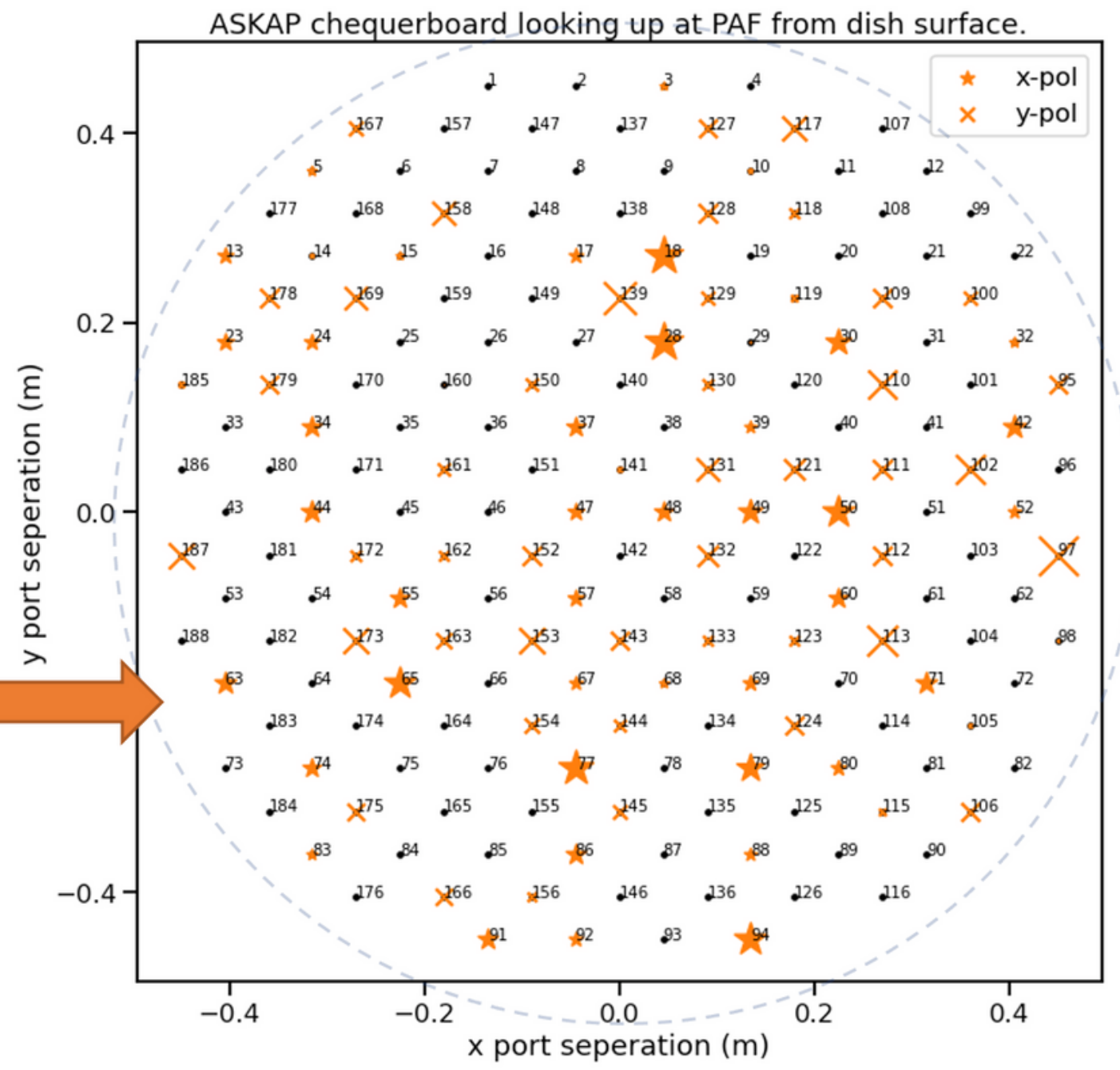
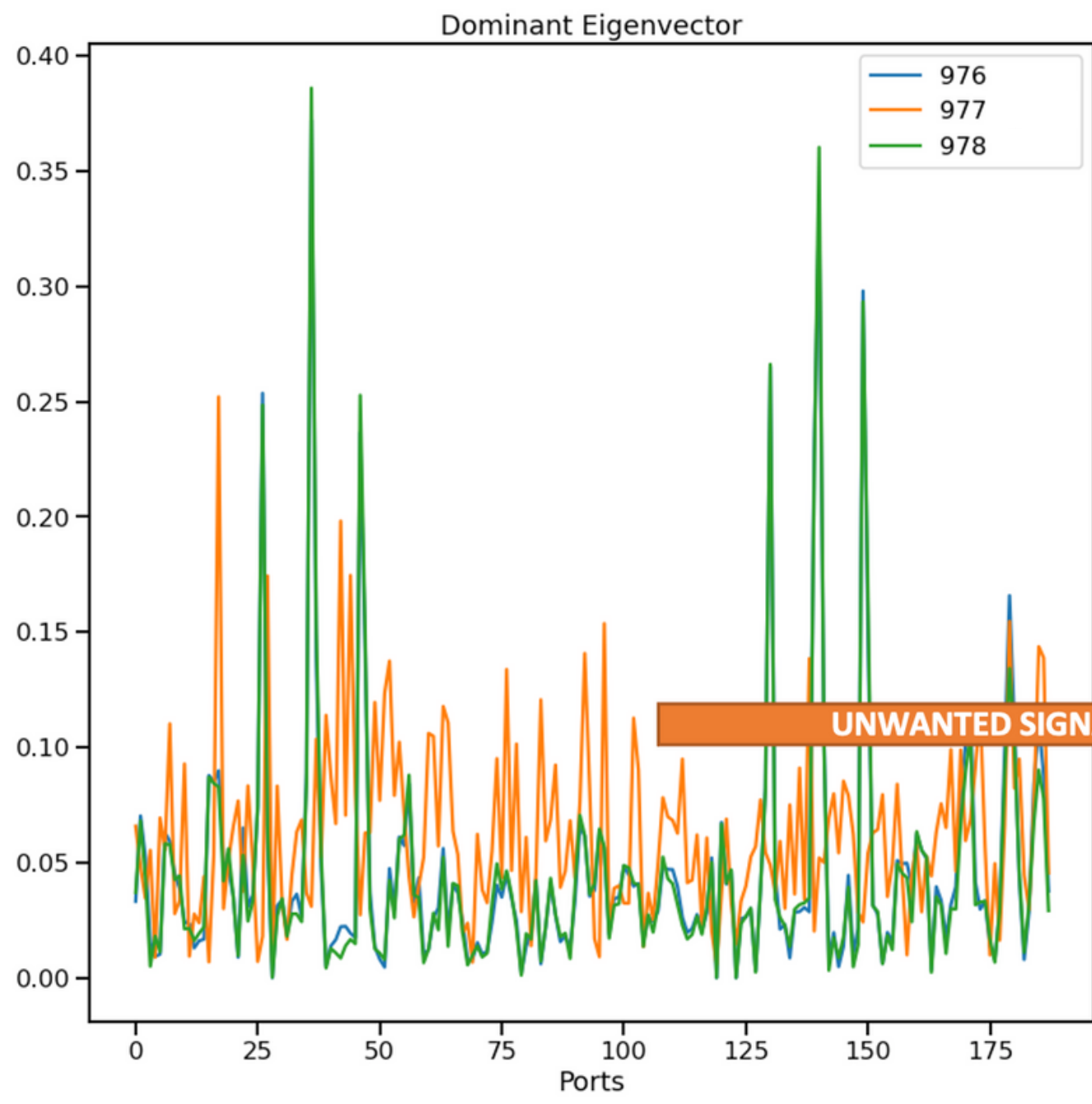
$$P_{obliq} = a_s (a_s^H P_{ortho} a_s)^{-1} a_s^H P_{ortho}$$



UNWANTED SIGNAL



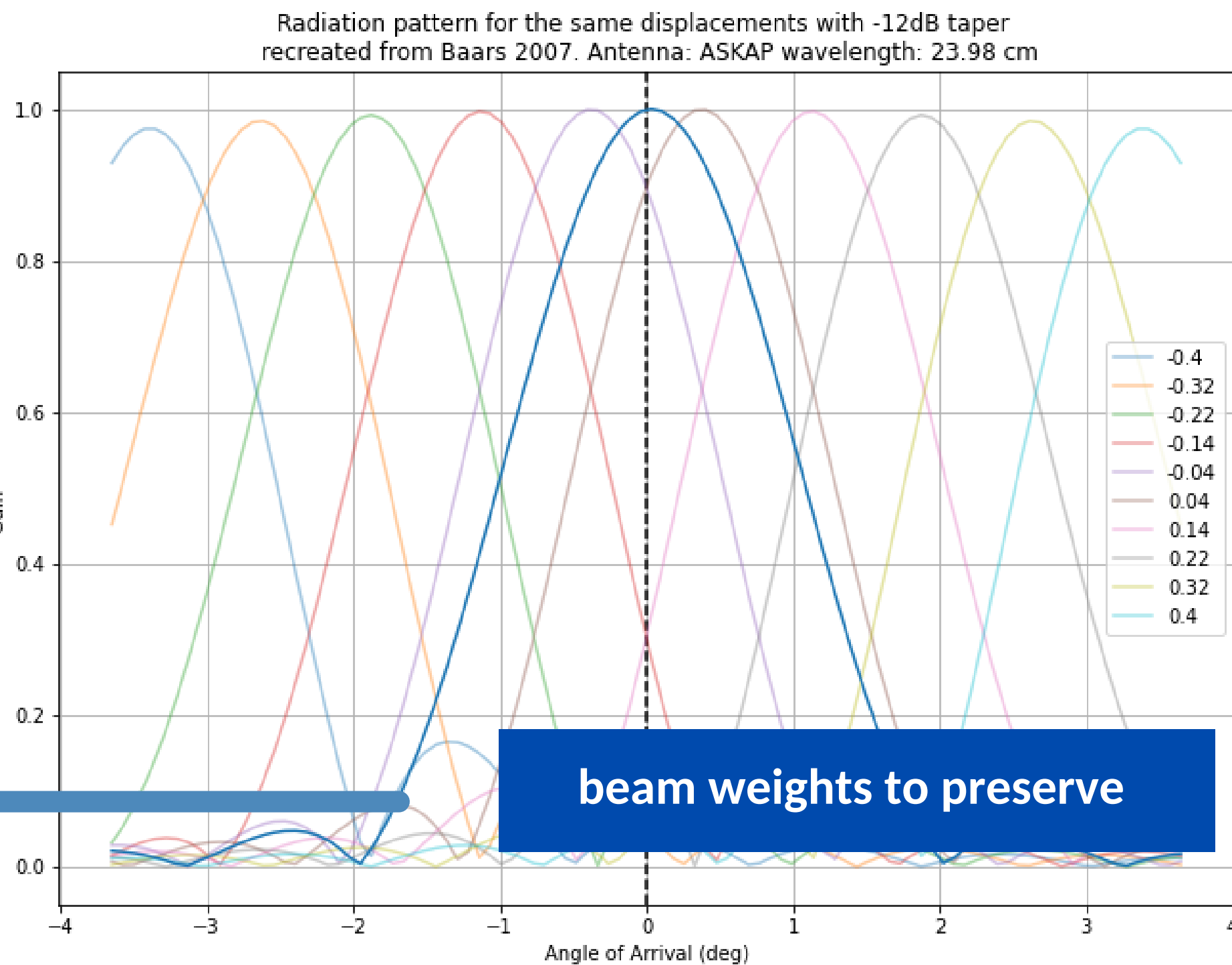
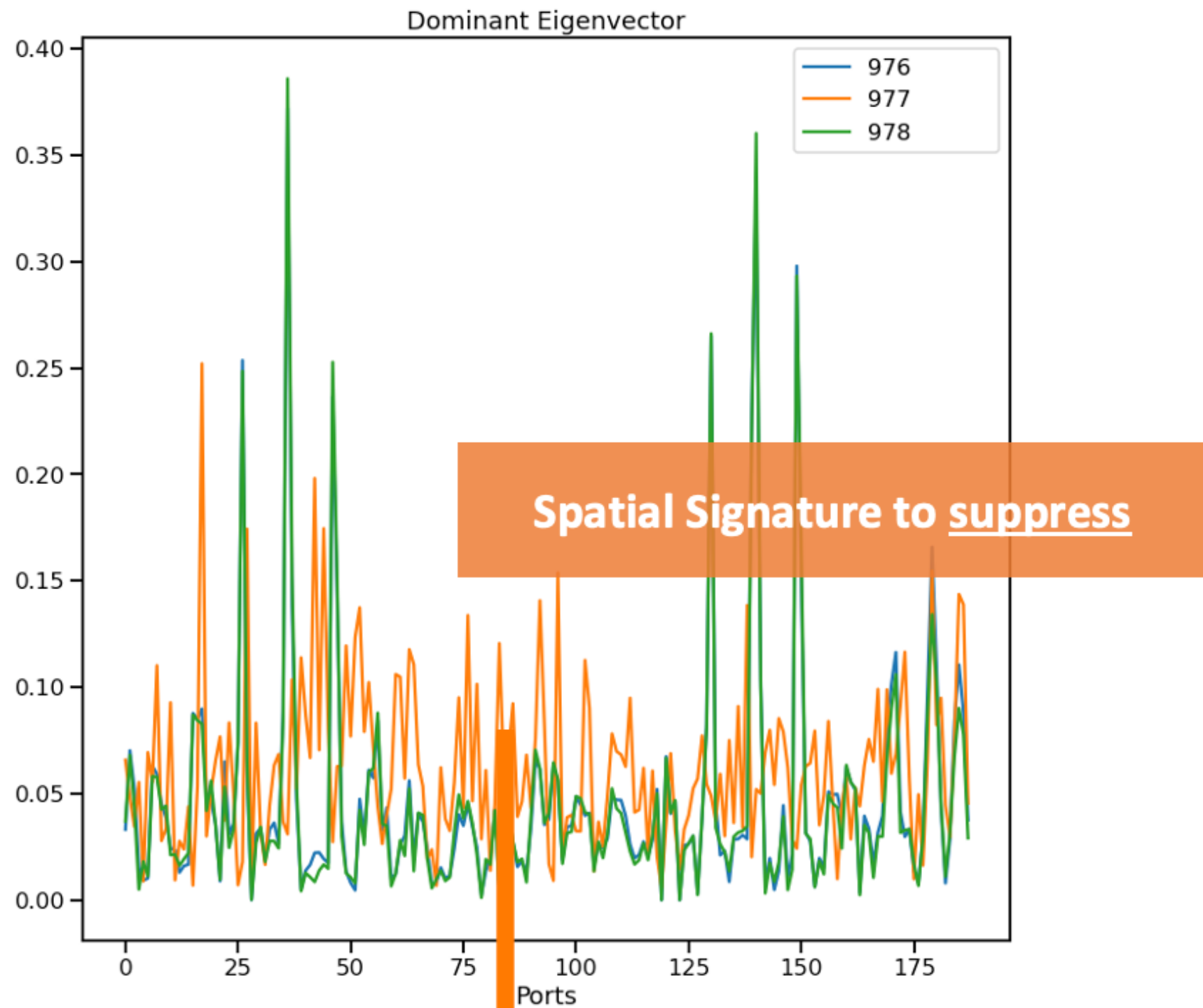
Eigenvalue decomposition to isolate the 'direction' of the interferer



Spatial signature to suppress

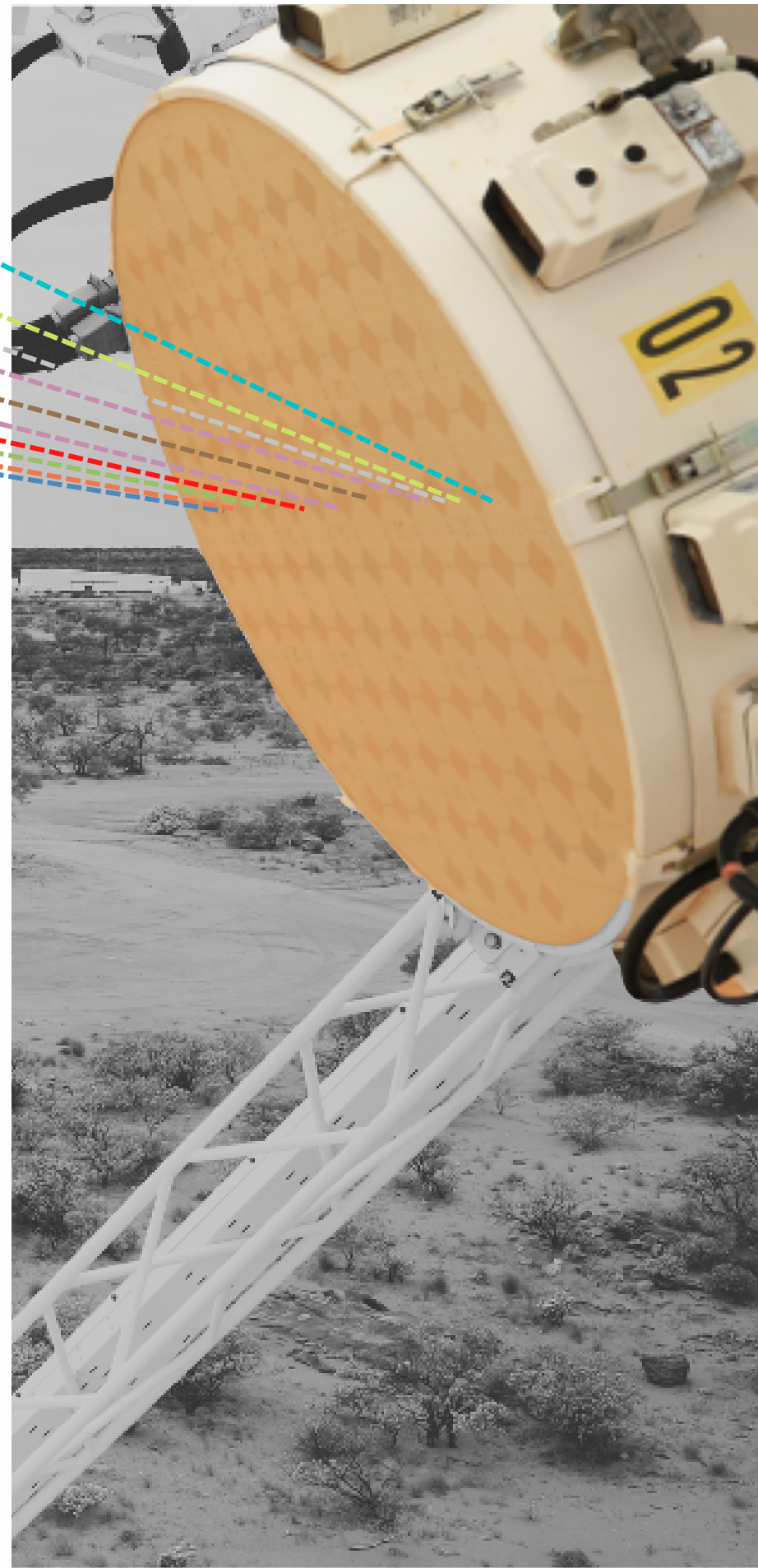
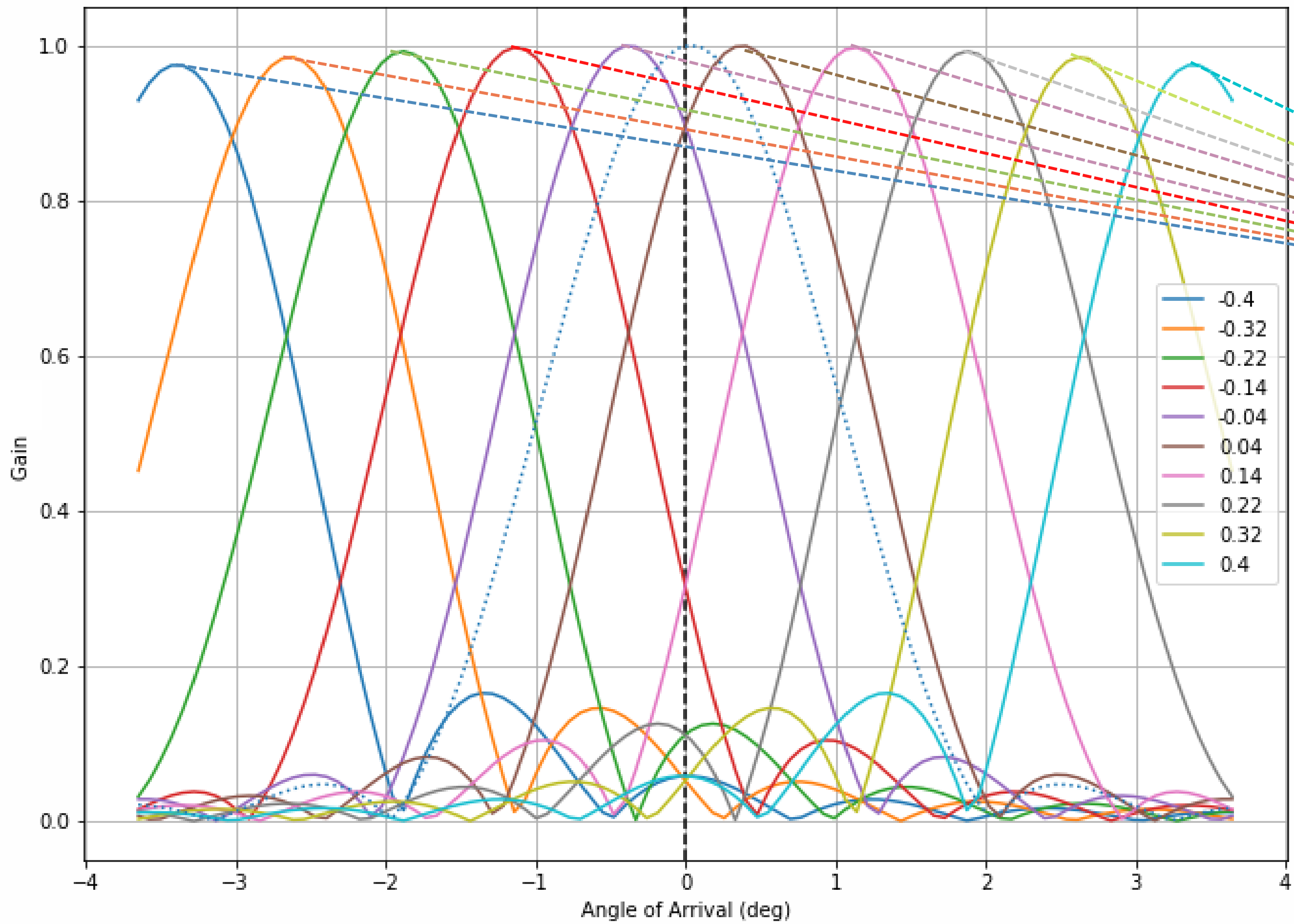


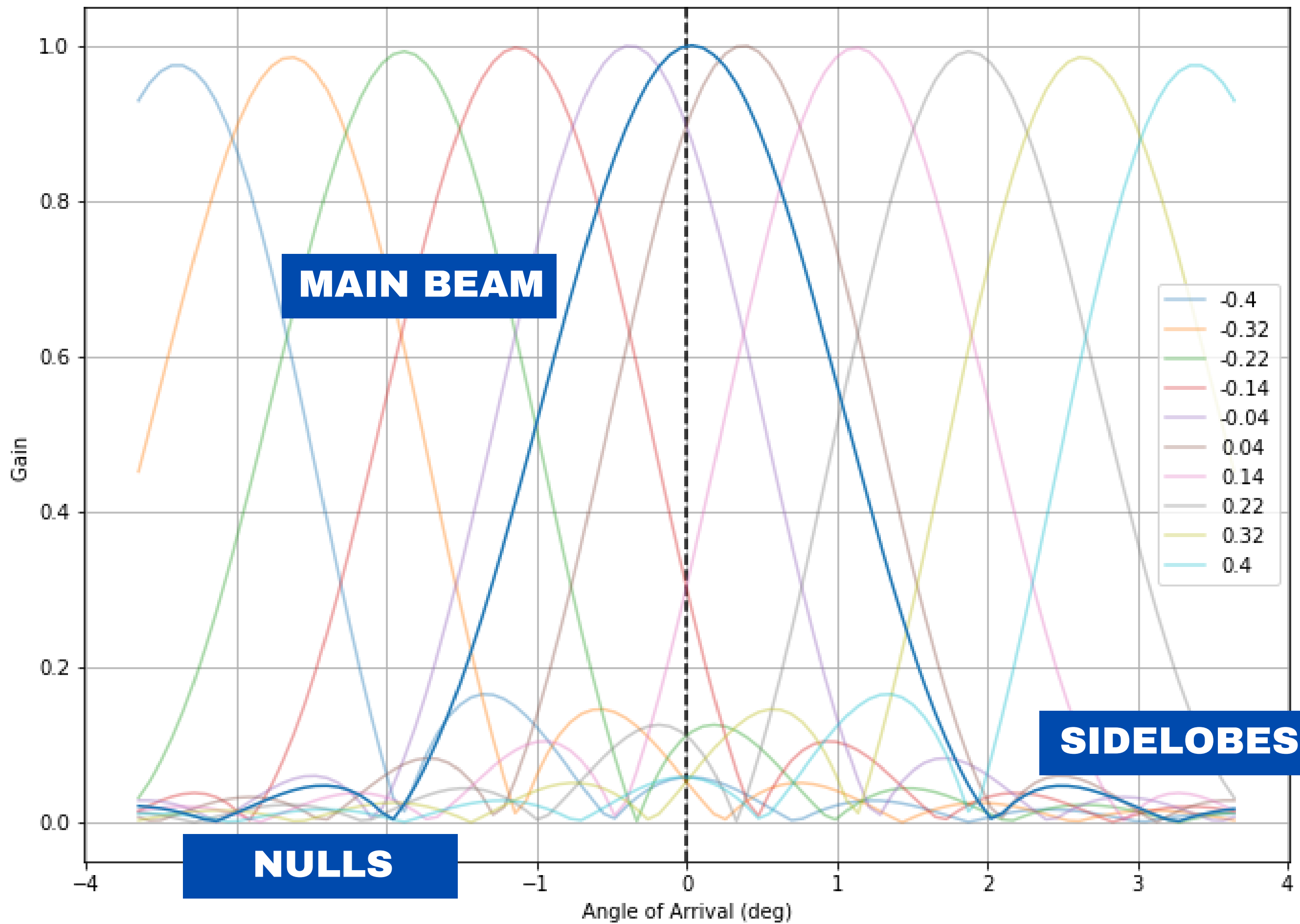
interference arrives at each port with different amplitude and phase
a unique spatial signature

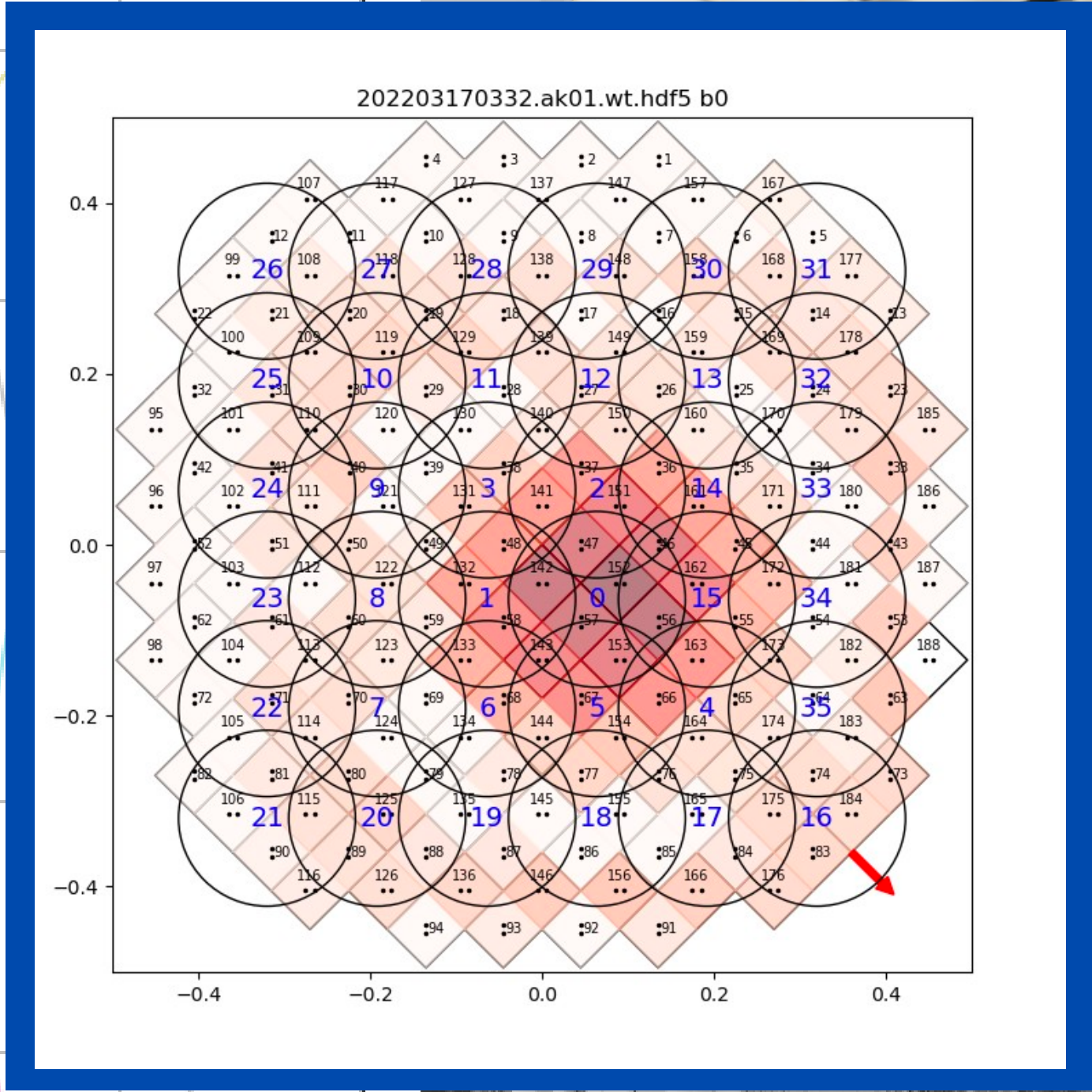
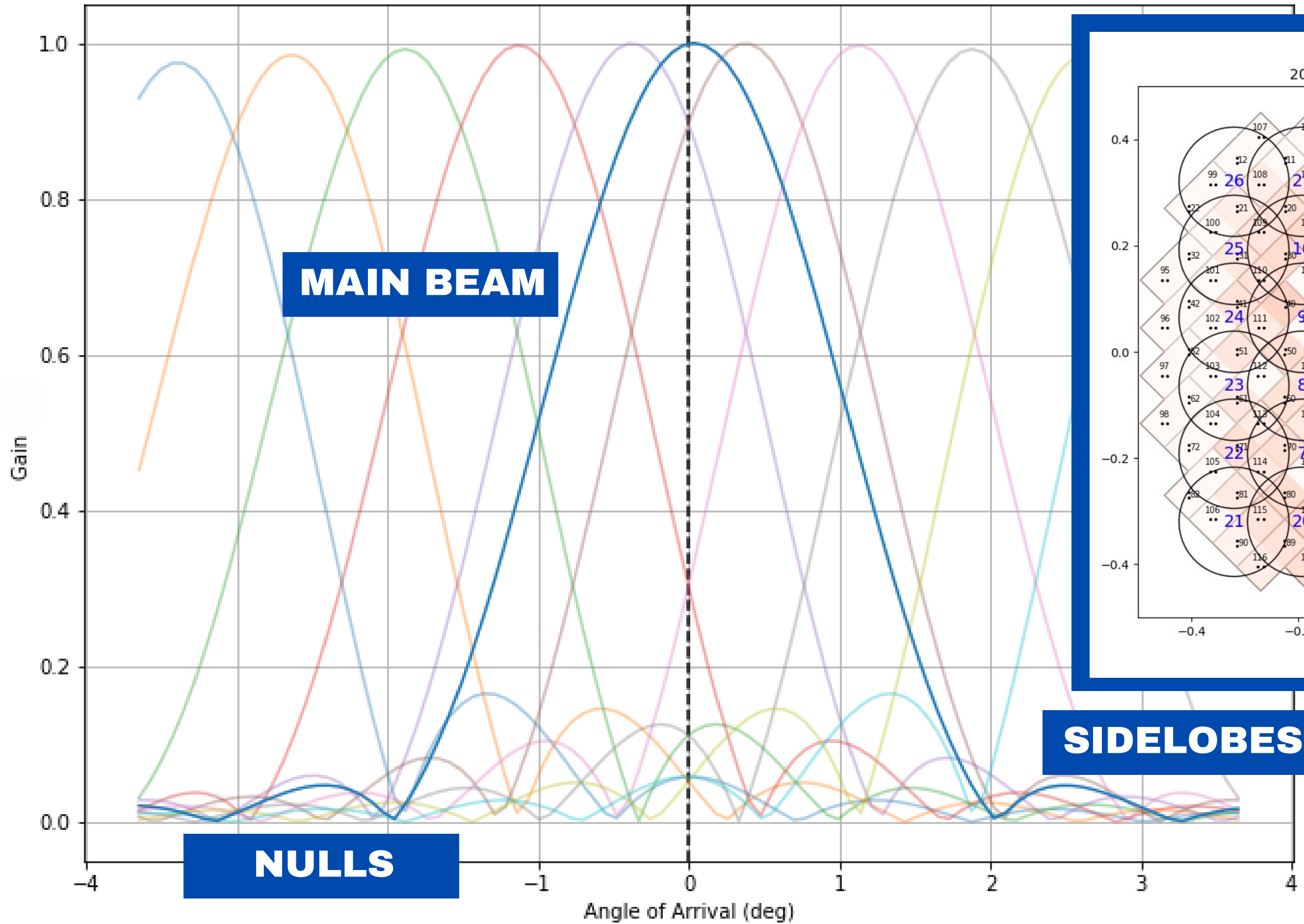


$$P_{ortho} = I - a_r (a_r^H a_r)^{-1} a_r^H$$

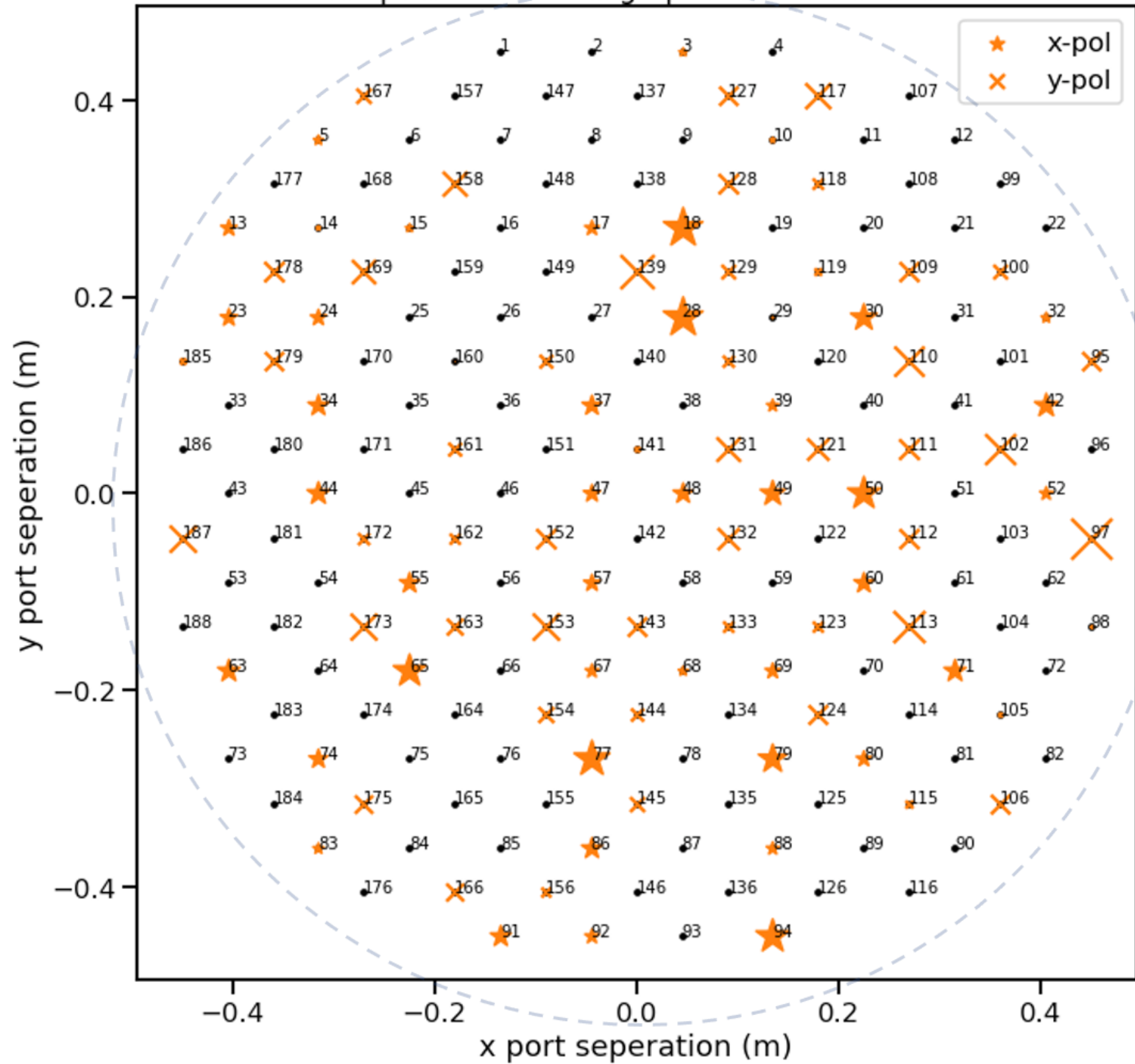
$$P_{obliq} = a_s (a_s^H P_{ortho} a_s)^{-1} a_s^H P_{ortho}$$





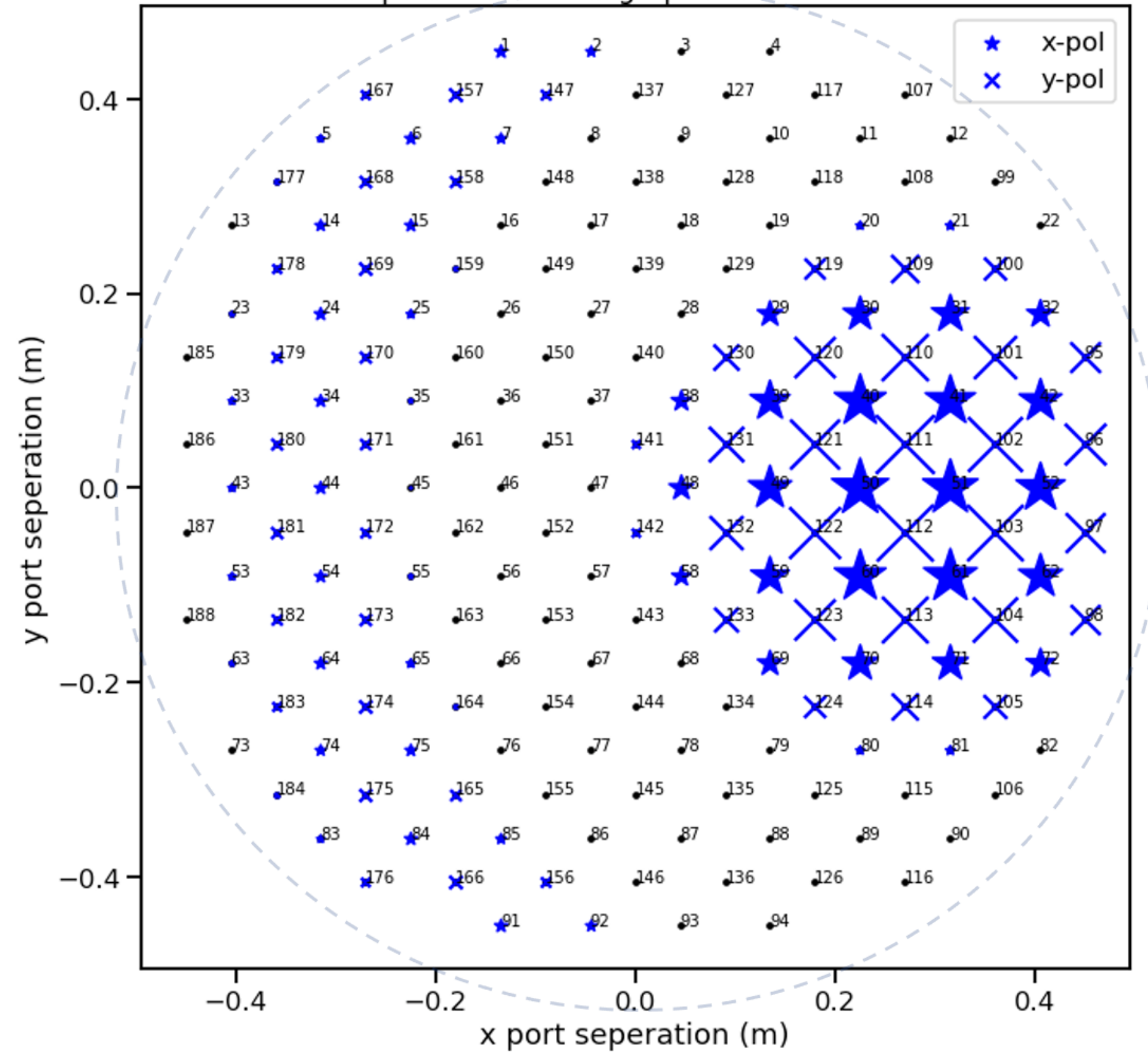


ASKAP chequerboard looking up at PAF from dish surface.



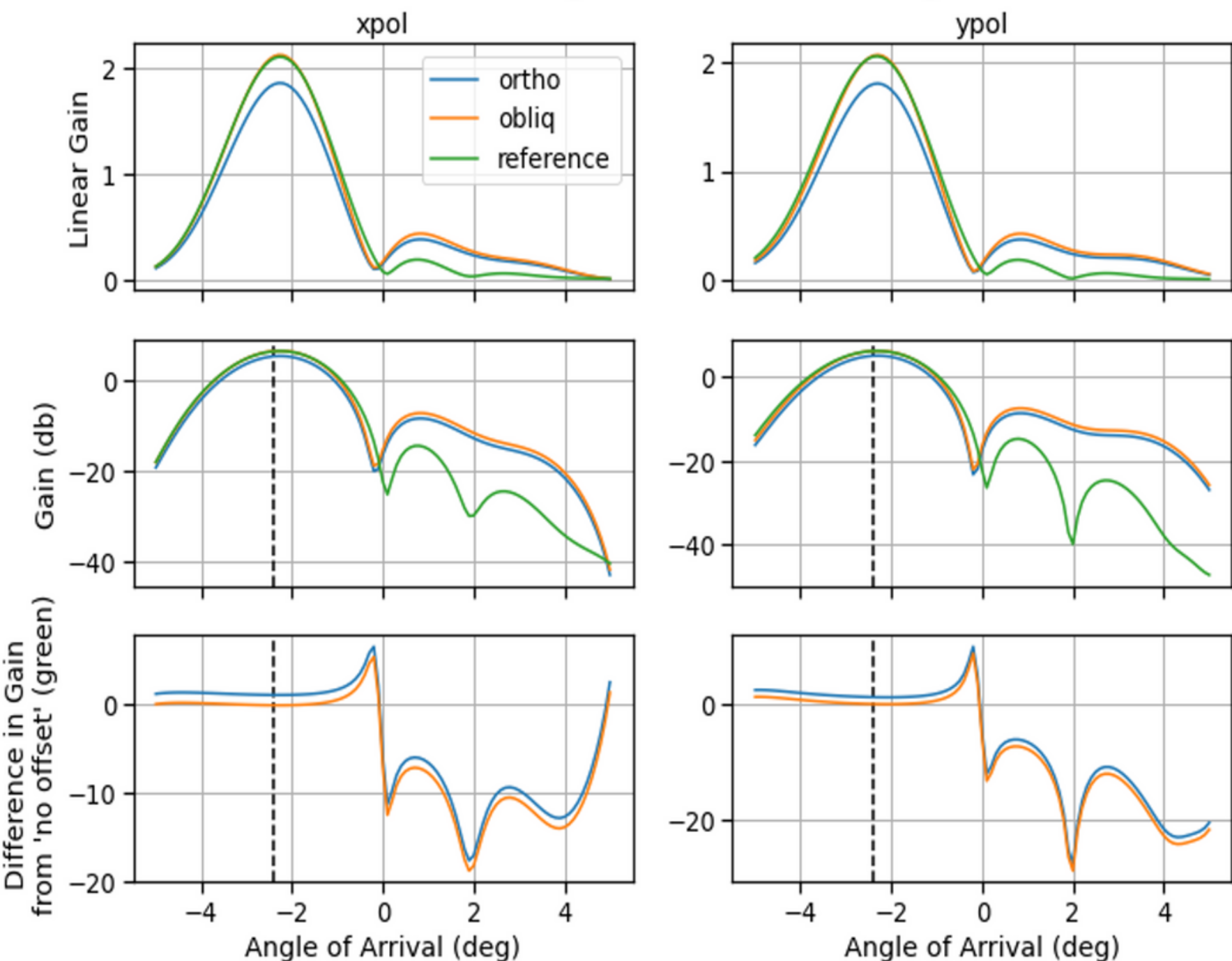
spatial signature to supress

ASKAP chequerboard looking up at PAF from dish surface.



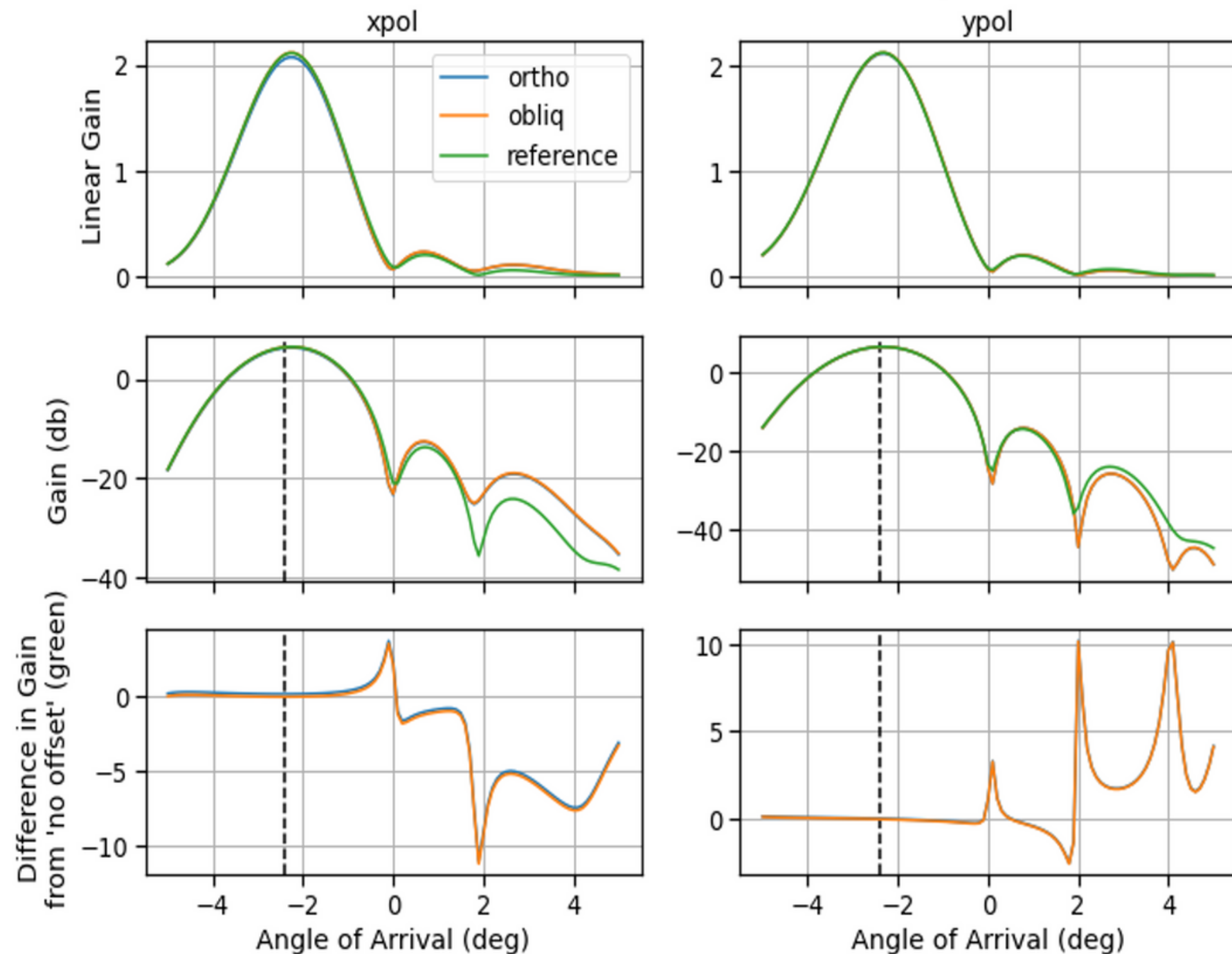
beamweights to preserve

Directivity Diagram of ASKAP 12m antenna at 976.44 MHz.
Beamforming in the direction -2.4 deg



Coherent Phase on all ports

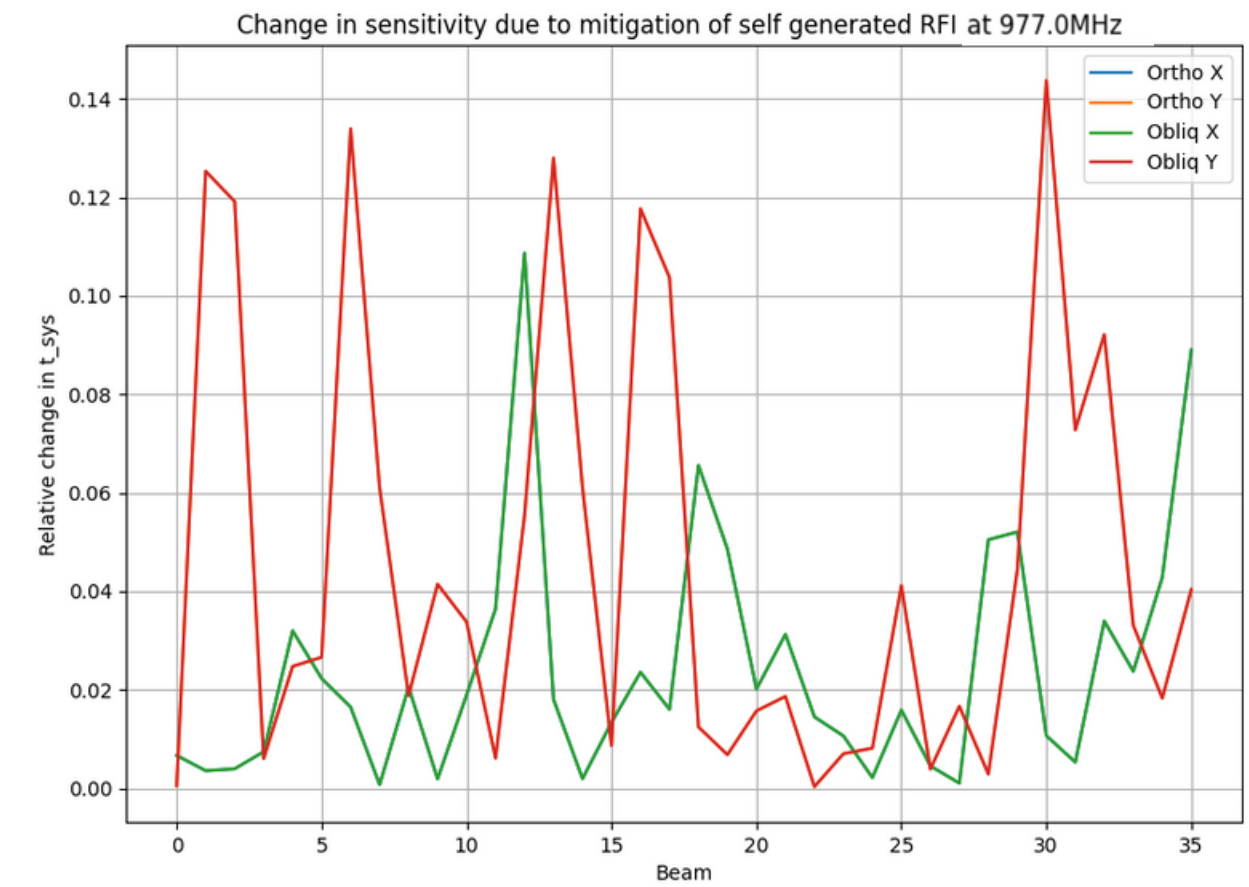
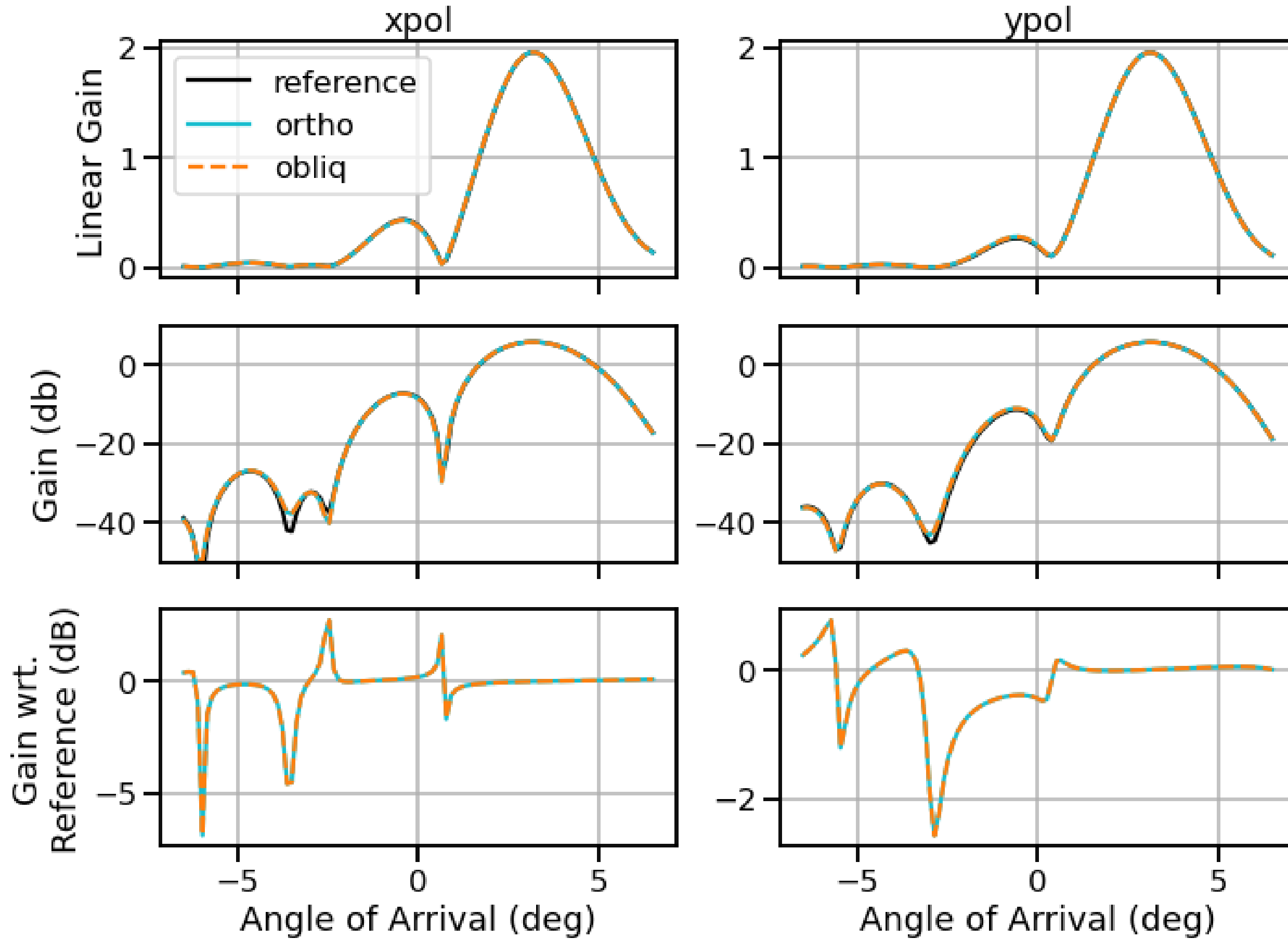
Directivity Diagram of ASKAP 12m antenna at 976.44 MHz.
Beamforming in the direction -2.4 deg



Different Phase on all ports

Variations to the beampatterns are reduced when the phase of the generated signal is randomized, particularly in the sidelobes.

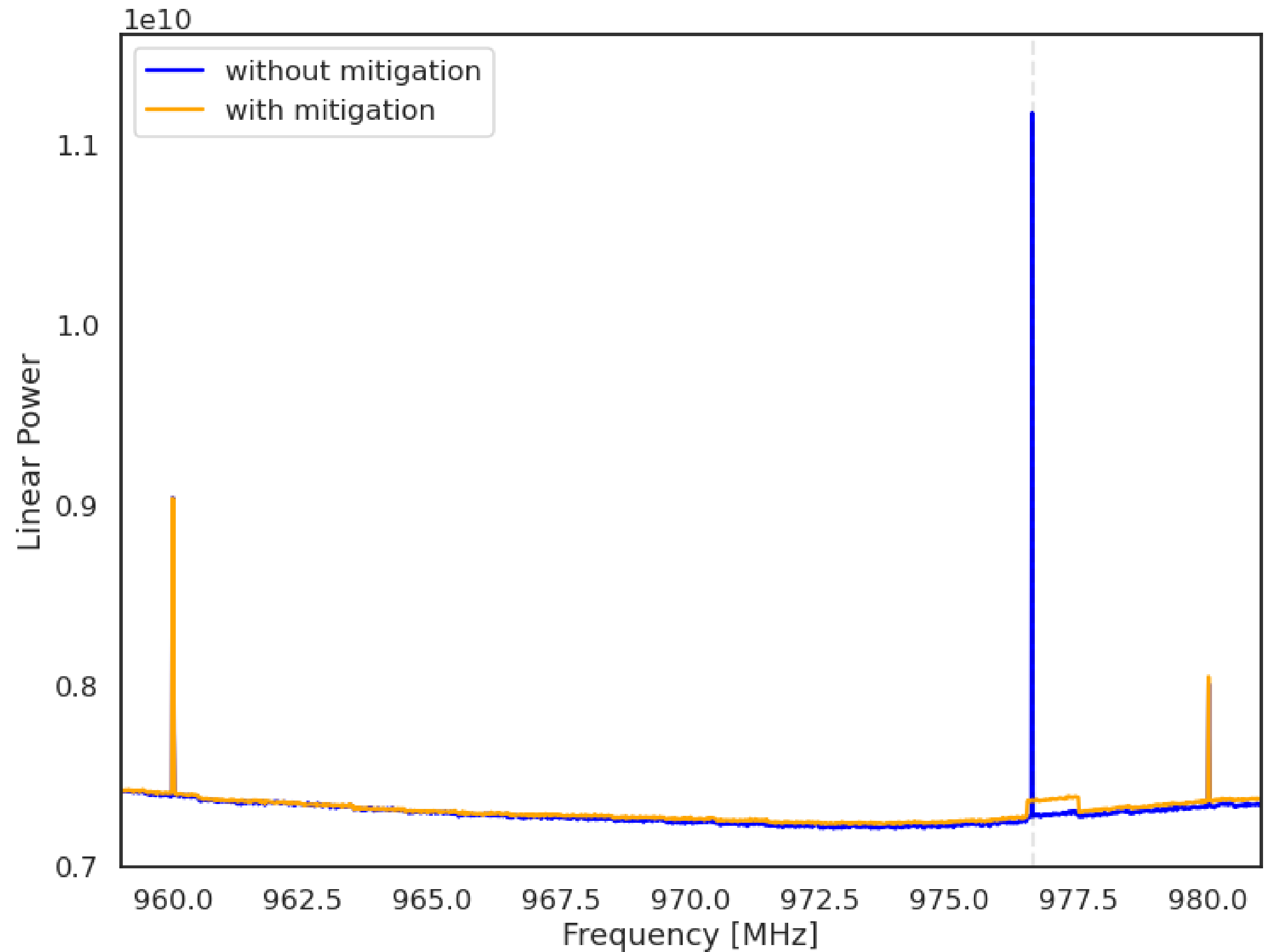
Directivity Diagram of ASKAP 12m antenna at 976.44 MHz. Beam26



SUPPRESSED OF
UNWANTED
NARROWBAND SIGNAL
BY OVER 31 DB, TO
THE NOISE FLOOR

REDUCTION IN
SENSITIVITY OF JUST
1.5 %.

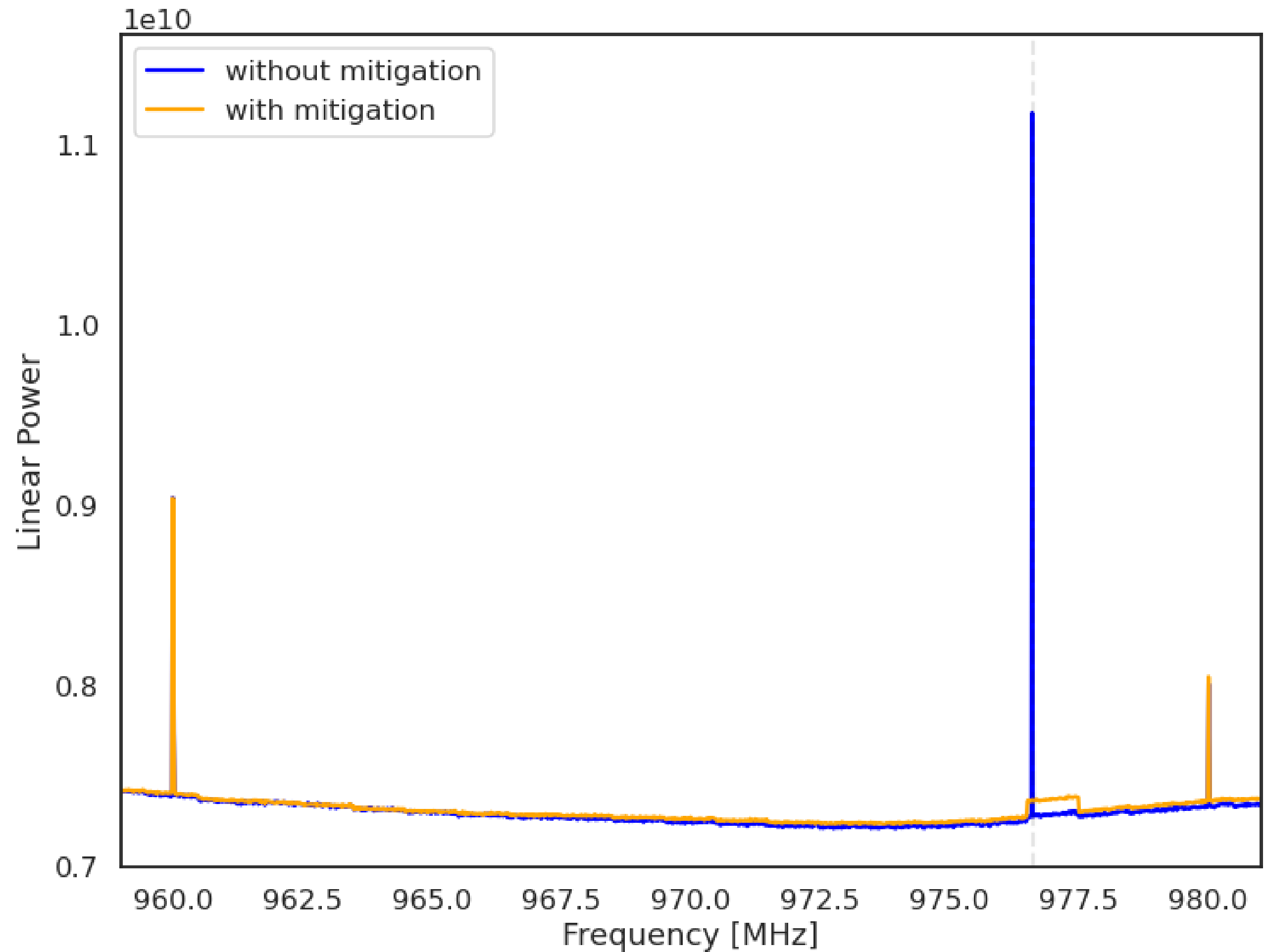
NO SIGNIFICANT
EXPECTED CHANGE TO
THE GAIN IN THE MAIN
BEAM AND FIRST
SIDELOBE.



GIVEN THIS KIND OF INFORMATION,

ASTRONOMERS ARE BETTER POSITIONED TO UNDERSTAND WHAT THE MITIGATION ALGORITHM IS DOING,

AND DECIDE ON THE USE OF THIS RFI MITIGATION TECHNIQUE BASED ON THEIR SCIENTIFIC GOALS.



Questions?

Next Steps:

- Measure sensitivity across the array and do holography of static use case on ASKAP.
 - Imaging and absorption/emission spectra, with and without mitigation.
- Model, implement and evaluate the dynamic case:
 - adjusting beam weights continuously throughout the observation.
 - shortlist suitable candidate RFI.
 - mitigate more pervasive RFI from moving sources, e.g. signals from satellites & aircraft.
 - starting with slower moving interferes
- Parkes CryoPAF

Liroy Lourenço

liroy.lourenco@sydney.edu.au

liroy.lourenco@csiro.au

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THE UNIVERSITY OF
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