OVERVIEW AND OUTLOOK OF SPACE SURVEILLANCE SYSTEMS AT FRAUNHOFER FHR

PAF WORKSHOP NOV. 2022

© Fraunhofer FHR

# Andreas Froehlich et. al

# Tasks of Space Missions



# SPACE DEBRIS – a growing problem

- 29.000 Objects > 10 cm, of which ~1.400 are active satellites
- $\sim$  75% of these objects are in LEO<br>(Height: 200 2.000 km) **E DEBRIS – a growing problem**<br> **C**esa<br>
0 Objects > 10 cm,<br>
ich ~1.400 are active satellites<br>
~ 75% of these objects are in LEO<br>
(Height: 200 – 2.000 km)<br>~ 9% of these objects are in GEO<br>
(Height 36.000 km)<br>
~ 00 Objects
	- $\blacksquare$  ~ 9% of these objects are in GEO (Height 36.000 km)
- $\blacksquare$  750.000 Objects > 1 cm only visible with high performance radar systems such as TIRA
- The number of space debris objects is increasing

 $\sim$  9% of these objects are in GEO<br>
(Height 36.000 km)<br>  $\sim$  750.000 Objects > 1 cm<br>
only visible with high performance radar systems<br>
such as TIRA<br>
The number of space debris objects is increasing<br>
steadily 1 lan 1980 1 lan 2000 1 Jan 2010 1 lan 1990 Reference Epoch

# Current space debris population

- All debris are human-made
- Incidents with numerous new debris particles:
	- 2007: A Chinese rocket intentionally destroyed a Chinese weather satellite
	- 2008: A US reconnaissance satellite was destroyed with an American rocket
	- 2009: Collision between a US communications satellite and a Russian satellite
- Horror scenario: Chain reaction ("Kessler Syndrome") 7.5 cm (Fraunhofer EMI)



Effect of impact of a 1-centimeter aluminum ball traveling at 6,500 m/s on a solid aluminum plate with a thickness of

# The space observation radar TIRA

TIRA supports all phases of space missions

- **Starts and first operation phase (LEOP)**
- Highly precise trajectory determination
- **Failure analysis**
- **Analysis of collision risks**
- **Monitoring of operations with robots**
- **Intrinsic rotation analysis**
- Supporting "Re-entry" and "De-orbiting" Manoeuvres





# The space observation radar TIRA

## Technical data

- Target tracking and imaging radar
- Most powerful system in Europe
	-
	-
	-
- **The space observation radar TIRA**<br> **Example 18 manufold act**<br> **Example 18 manufold ma** power), sensitivity: Detection of objects > 2 cm in 1000 km (with Effelsberg: < 1 cm)
- Target imaging radar: Ku-band (16.7 GHz, resolution better than 20



# Analysis of damage and proper motion: ADEOS I



- **Total failure of power supply through** demolition of solar panel (24m x 3m x 0.5mm)
- Radar imaging is the only feasible damage assessment possibility!







# Detection and Cataloguing of Space Debris **Detection and Cataloguing of Space Debris**<br>GESTRA – German Space Surveillance Radar<br>■ Federal government outlines necessity of establishing a national

- competence center for documentation and evaluation of the current situation in space
- Quasi-monostatic pulsed phased array radar in L-band  $(^{\sim} 1.3$  GHz)
- Surveillance in orbital heights of 300-3000 km
- Electronic swiveling through HF-based modification of the wave front during transmission or digital beam forming during reception
- Numerous operating modes (surveillance volumes vs. detection



# Detection and Cataloguing of Space Debris **Detection and Cataloguing of Space Debris<br>GESTRA – German Space Surveillance Radar<br>■ GESTRA consists of:**<br>■ 2 containary and receiving (BY) and and **Exection and Cataloguing of Space Debris**<br> **EXPRA – German Space Surveillance Radar**<br> **EXPRA consists of:**<br>
■ 2 containers - one receiving (RX) and one<br>
transmitting (TX) unit<br>
■ both subsystems have independent infrastr

# GESTRA consists of:

- transmitting (TX) unit
- $\blacksquare$  both subsystems have independent infrastructure (energy, cooling, climate control, etc.)
- size each: 18 m x 4 m x 4 m; weight about 90 t
- each unit (RX&TX) contains:
	- a phased array antenna with 256 individual elements
	- mounted on top of a 3-D positioner



# Further detection concepts and projects GESTRA EU-SST & GESTRA networks

- the energy reflected away can be collected by a bi-static sensor configuration as implemented by EUSST (lower graph)
- alternatively, larger search volumes can be "scanned" simultaneously
- more accurate tracks can be calculated by combining multiple sensor contributions **STRA EU-SST & GESTRA networks**<br>the energy reflected away can be collected by a bi-static sensor<br>configuration as implemented by EUSST (lower graph)<br>alternatively, larger search volumes can be "scanned"<br>simultaneously<br>more
- the sensitivity of the overall system is significantly increased by<br>bi- or even multi-static arrangement of the sensors
- high-precision synchronization and appropriate signal processing are required for implementation





# Further detection concepts and projects GESTRA TX2





# Further detection concepts and projects GESTRA network configurations

Improvement of the minimum detectable RCS in dB at 1,000 km altitude for different GESTRA network configurations.

- C1: GESTRA
- **C2: GESTRA + GESTRA EUSST**
- C3: GESTRA + GESTRA EUSST / GESTRA TX2
- C4: GESTRA + GESTRA EUSST + GESTRA TX2
- Attention: RCS values are normalized to the minimum detectable RCS of GESTRA (0 dB)
- **Results:** 
	- C2: Detection power increases by up to 2.3 dB
	- C3/C4: Detection power increases by up to 6,3 dB<br>(c) Configuration C3
- Detection of targets that are up to four times smaller



# Cryo-Cooled 37-Element Phased Array Radar Receiver



## Components:

- 
- **Receiver**<br>
Components:<br>
1 RF-Unit-Cell<br>
2 2-Stage Cryocooler<br>
3 Vacuum Vessel **Receiver**<br>2 2-Stage Cryocooler<br>2 2-Stage Cryocooler<br>2 Vacuum Vessel **Receiver**<br>
Components:<br>
1 RF-Unit-Cell<br>
2 2-Stage Cryocooler<br>
3 Vacuum Vessel<br>
Diameter 1.5 m
- 

## Diameter 1.5 m

# Cryo-Cooled 37-Element Phased Array Radar Receiver **dar Receiver**<br>Vaccum Vessel with Cryocooler<br>1 Sumitomo Cryocooler<br>2 Vacuum Vessel **dar Receiver**<br>Vaccum Vessel with Cryocooler<br>1 Sumitomo Cryocooler<br>2 Vacuum Vessel<br>3 50K-Stage **dar Receiver**<br>Vaccum Vessel with Cryocooler<br>1 Sumitomo Cryocooler<br>2 Vacuum Vessel<br>3 SOK-Stage<br>4 4K-Stage **dar Receiver**<br>
Vaccum Vessel with Cryocooler<br>
1 Sumitomo Cryocooler<br>
2 Vacuum Vessel<br>
3 50K-Stage<br>
4 4K-Stage<br>
5 Copper-Shield **dar Receiver**<br>Vaccum Vessel with Cryocooler<br>1 Sumitomo Cryocooler<br>2 Vacuum Vessel<br>3 SOK-Stage<br>4 4K-Stage<br>5 Copper-Shield<br>6 Vacuum Barrier **dar Receiver**<br>Vaccum Vessel with Cryocooler<br>1 Sumitomo Cryocooler<br>2 Vacuum Vessel<br>3 50K-Stage<br>4 4K-Stage<br>5 Copper-Shield<br>6 Vacuum Barrier<br>7 Copper Rod **Samillar Receiver**<br>
Vaccum Vessel with Cryocooler<br>
1 Sumitomo Cryocooler<br>
2 Vacuum Vessel<br>
3 SOK-Stage<br>
4 4K-Stage<br>
5 Copper-Shield<br>
6 Vacuum Barrier<br>
7 Copper Rod<br>
8 Radom



Vaccum Vessel with Cryocooler<br>1 Sumitomo Cryocooler<br>2 Vacuum Vessel<br>3 SOK-Stage<br>4 4K-Stage<br>5 Copper-Shield<br>6 Vacuum Barrier<br>7 Copper Rod<br>8 Radom<br>9 Vacuum Pump Interface Vaccum Vessel with Cryocooler<br>1 Sumitomo Cryocooler<br>2 Vacuum Vessel<br>3 SOK-Stage<br>4 4K-Stage<br>5 Copper-Shield<br>6 Vacuum Barrier<br>7 Copper Rod<br>8 Radom<br>9 Vacuum Pump Interface<br>10 Lower Flange

- 
- 
- 
- 
- 5 Copper-Shield
- 
- 
- 
- Vacuum Vesser Minter, yeelooner<br>
1 Sumitomo Cryocooler<br>
2 Vacuum Vessel<br>
3 50K-Stage<br>
4 4K-Stage<br>
5 Copper-Shield<br>
6 Vacuum Barrier<br>
7 Copper Rod<br>
8 Radom<br>
9 Vacuum Pump Interface<br>
10 Lower Flange 1 Sumitomo Cryocooler<br>
2 Vacuum Vessel<br>
3 50K-Stage<br>
4 4K-Stage<br>
5 Copper-Shield<br>
6 Vacuum Barrier<br>
7 Copper Rod<br>
8 Radom<br>
9 Vacuum Pump Interface<br>
10 Lower Flange
- 

# Some simulation results

Maximum deformation along the Z axis was 10.45 mm

-> To reduce this value, either the thickness of the vacuum vessel can be increased or the 4 stand feet can be scaled higher

Temperature distribution in the LNA area was between 23.5 and 29.7 K and in the antenna area between 108.4 and 134.2 K -> Temperature differences can be reduced by increasing the cross-sectional area

 $\triangleright$  Based on the simulation results, optimizations can be made and incorporated into the hardware structure

A. Froehlich et al 2022 IOP Conf. Ser.: Mater. Sci. Eng. 1240 012102

# Conclusion

- Showed Space Debris is a rising problem for our infrastructures and services
- TIRA allows to support all stages of space missions
- GESTRA family a novel phased array systems for detection and catalogue space debris
- New concepts of GESTRA systems such as GESTRA Networks or cryocooled receivers improve the performance further





# Thank you for your attention

# Thanks to all colleagues from Fraunofer

Contact:

Andreas Froehlich

andreas.froehlich@fhr.Fraunhofer.de





# Cryo-Cooled 37-Element Phased Array Radar Receiver **Radar Receiver<br>
RF unit cell:<br>
1 50K-Plate<br>
2 4K-Plate Radar Receiver<br>
RF unit cell:<br>
1 50K-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables Radar Receiver<br>
RF unit cell:**<br>
1 50K-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer **READ FRECEIVER<br>
2 AK-Plate<br>
2 AK-Plate<br>
2 AK-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer<br>
5 LNA (Low noise amplifier) 7 Radar Receiver**<br> **RF unit cell:**<br>
1 50K-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer<br>
5 LNA (Low noise amplifier)<br>
6 LNA Fixing Plate (Copper) **READ TRECEIVET<br>
READ FORMALE:**<br>
1 SOK-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer<br>
5 LNA (Low noise amplifier)<br>
6 LNA Fixing Plate (Copper)<br>
7 Copper Stands **7 Radar Receiver**<br> **RF unit cell:**<br>
1 50K-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer<br>
5 LNA (Low noise amplifier)<br>
6 LNA Fixing Plate (Copper)<br>
7 Copper Stands<br>
8 Patch-Antenna



- 
- 
- 
- 
- RF unit cell:<br>
1 SOK-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer<br>
5 LNA (Low noise amplifier)<br>
6 LNA Fixing Plate (Copper)<br>
7 Copper Stands<br>
8 Patch-Antenna<br>
9 Cavity RF unit cell:<br>1 50K-Plate<br>2 4K-Plate<br>3 Cooper Rod for RF-Cables<br>4 Copper Spacer<br>5 LNA (Low noise amplifier)<br>6 LNA Fixing Plate (Copper)<br>7 Copper Stands<br>8 Patch-Antenna<br>9 Cavity<br>10 GFK-Spacer RF unit cell:<br>
1 50K-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer<br>
5 LNA (Low noise amplifier)<br>
6 LNA Fixing Plate (Copper)<br>
7 Copper Stands<br>
8 Patch-Antenna<br>
9 Cavity<br>
10 GFK-Spacer 1 50K-Plate<br>
2 4K-Plate<br>
3 Cooper Rod for RF-Cables<br>
4 Copper Spacer<br>
5 LNA (Low noise amplifier)<br>
6 LNA Fixing Plate (Copper)<br>
7 Copper Stands<br>
8 Patch-Antenna<br>
9 Cavity<br>
10 GFK-Spacer
- 
- 
- 
- 
- 

# Preparing and Optimizing of 3D-Simulation Model





# Simulation Model and Project Scheme<br>
Intermal Model<br>
Streat flow of the cables

# ■ Thermal Model

• Heat flow of the cables



• Heat transfer through thermal radiation and the state of the structure that the structure fact to the structure structur





# Simulation Model and Project Scheme<br>
I Mechanical Model

# **Mechanical Model**



- Fixed storage on the 4 feet of the vacuum vessel
- Atmospheric pressure in external surfaces



# **Simulation Model and Project Scheme**<br>
Modal Analysis<br> **Determination of the eigenfrequencies by the eigenvalue equation**<br>
Harmonic Analysis<br> **Response of the system at the determined natural frequencies**<br> **Response of th Simulation Model and Project Scheme**<br>
Modal Analysis<br> **In Determination of the eigenfrequencies by the eigenvalue equation**<br>
Harmonic Analysis<br> **Exponse of the system at the determined natural frequencies**<br> **Exponse Spec Simulation Model and Project Scheme**<br>
Modal Analysis<br> **A** Determination of the eigenfrequencies by the eigenvalue equation<br>  $(-\omega_1^2 \{M\} + [K])$ <br>
Harmonic Analysis<br> **A** Response of the system at the determined natural freq Simulation Model and Project Scheme<br>
<sub>odal Analysis</sub><br>
Determination of the eigenfrequencies by the eigenvalue equation

Modal Analysis

Harmonic Analysis

- $\blacksquare$  Response of the system at the determined natural frequencies
- 

Response Spectrum

- Determination of the deformation for a given response spectrum
- At a maximum angular velocity of  $5 \degree$ /s
- Determination of the eigenfrequencies by the eigenvalue equation<br>
(-ω<br>
Harmonic Analysis<br>
 Response of the system at the determined natural frequencies<br>
 Newton stimulating force<br>
Response Spectrum<br>
→ Determination o e system at the determined natural frequencies<br>
ulating force<br>
of the deformation for a given response spectrum<br>
angular velocity of 5 °/s<br>
ngular acceleration of 5 °/s2 results in a total acceleration of<br>  $a_{ges} = \sqrt{a_r^2 + a$

$$
a_{ges} = \sqrt{a_r^2 + a^2_t} = 61.31 \frac{mm}{s^2} \qquad .
$$

 $(-\omega_i^2 \quad [M] + [K]$ 



- -

$$
PSD = \frac{a^2_{\text{ges}}}{\Delta f} = 21.57 \text{ mm}^2 \text{s}^{-4} / Hz
$$



## Mechanics Model





# Mechanical Model

# Deformation



Deviation due to feet fixation on the vacuum vessel



# **Result and Evaluation<br>hermal Simulation<br>hermal Simulation**

# Thermal Simulation





# Thermal Simulation





# Grafische Elemente

# Kästen, Dfeile, Verbindungen und Linien (Auswahl) ! DIESE FOLIE AUS FINALER PRÄSENTATION LÖSCHEN !

 folgende Elemente können hier per Rechtsklick kopiert und an gewünschter Stelle in der neuen Präsentation per Rechtsklick wieder eingesetzt werden:



# Farben

# ! DIESE FOLIE AUS FINALER PRÄSENTATION LÖSCHEN !

- folgende Farben können über die Powerpoint-Farbauswahl hier aufgenommen und damit in der neuen Präsentation angewendet werden:
- Überschriften / Fließtext / Quellenangaben / Bildunterschriften / Grafikauszeichnungen
- Grafikauszeichnungen
- Aufzählungen / Nummerierungen erster Ebene / grafische Elemente
- Grafiken

# Fonds hinter Grafiken

