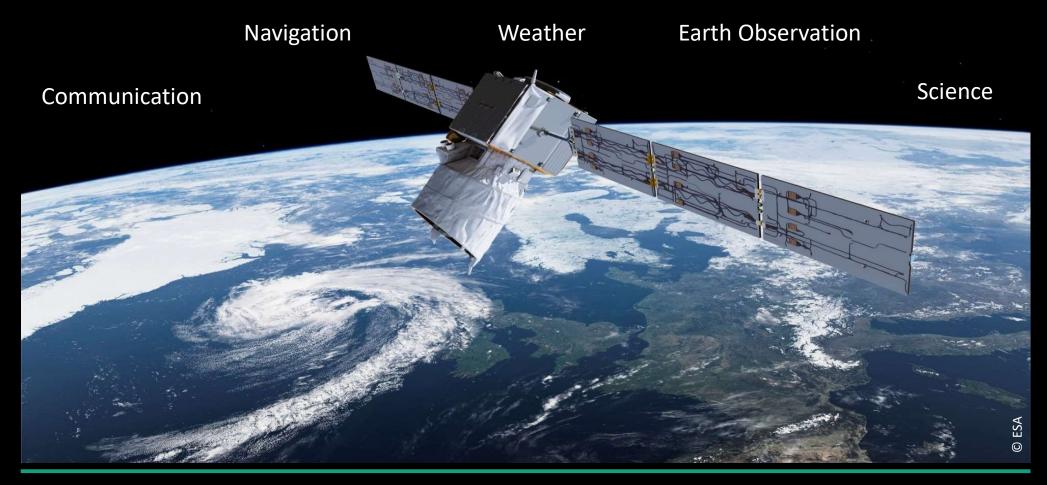


Tasks of Space Missions



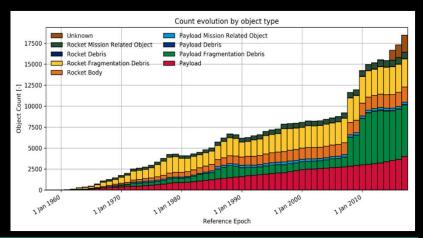


© ESA/DLR

SPACE DEBRIS – a growing problem

- 29.000 Objects > 10 cm, of which ~1.400 are active satellites
 - ~ 75% of these objects are in LEO (Height: 200 2.000 km)
 - ~ 9% of these objects are in GEO (Height 36.000 km)
- 750.000 Objects > 1 cm only visible with high performance radar systems such as TIRA
- The number of space debris objects is increasing steadily







Current space debris population

- All debris are human-made
- Incidents with numerous new debris particles:
 - 2007: A Chinese rocket intentionally destroyed a Chinese weather satellite
 - 2008: A US reconnaissance satellite was destroyed with an American rocket
 - 2009: Collision between a US communications satellite and a Russian satellite
- Horror scenario: Chain reaction ("Kessler Syndrome")

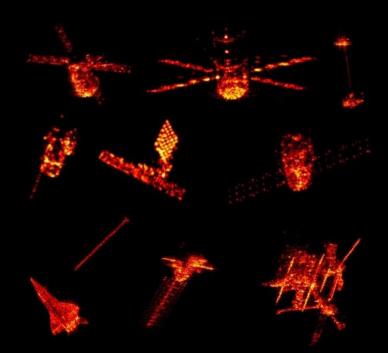


Effect of impact of a 1-centimeter aluminum ball traveling at 6,500 m/s on a solid aluminum plate with a thickness of 7.5 cm (Fraunhofer EMI)



The space observation radar TIRA

- TIRA supports all phases of space missions
 - Starts and first operation phase (LEOP)
 - Highly precise trajectory determination
 - failure analysis
 - Analysis of collision risks
 - Monitoring of operations with robots
 - Intrinsic rotation analysis
 - Supporting "Re-entry" and "De-orbiting" Manoeuvres





The space observation radar TIRA

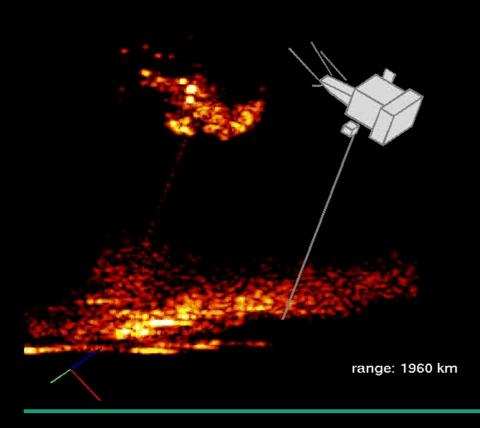
Technical data

- Target tracking and imaging radar
- Most powerful system in Europe
 - Positioning accuracy: 0.000172° 3 mm / 1000 m
 - Diameter of reflector: 34 m (radome 47.5 m)
 - Antenna speed 24° / s (360° in 15 s)
- Target tracking radar: L-band (1.333 GHz, up to 1.5 MW transmit power), sensitivity: Detection of objects > 2 cm in 1000 km (with Effelsberg: < 1 cm)
- Target imaging radar: Ku-band (16.7 GHz, resolution better than 20 cm)



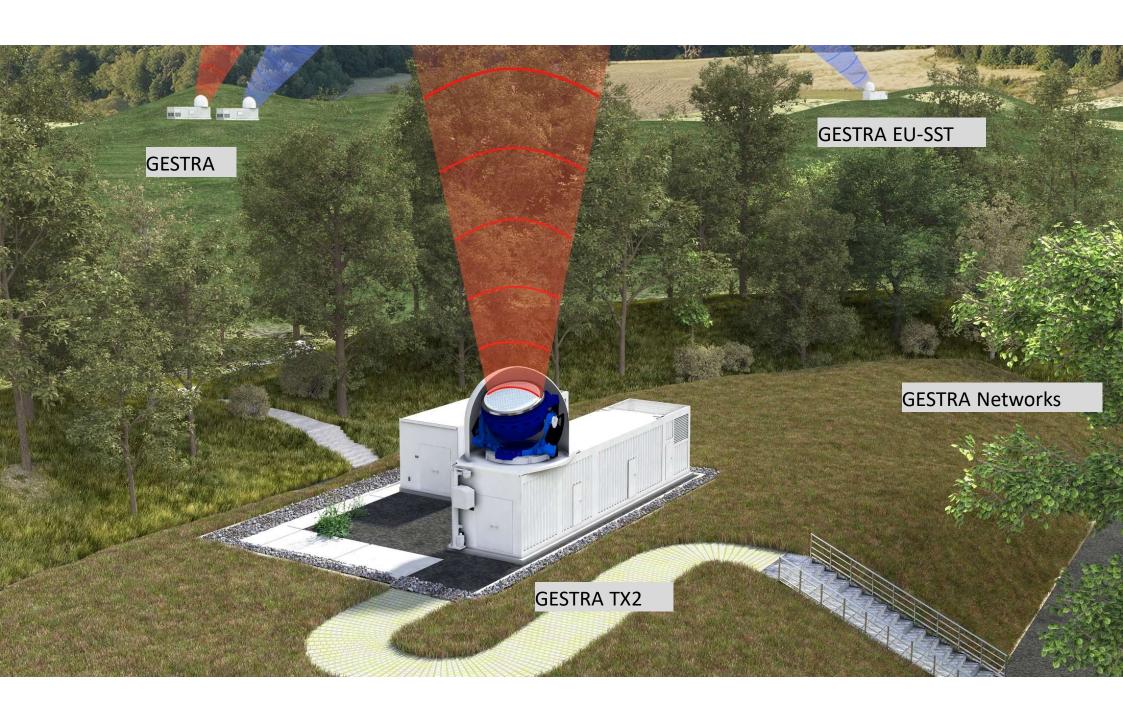


Analysis of damage and proper motion: ADEOS I



- Total failure of power supply through demolition of solar panel (24m x 3m x 0.5mm)
- Radar imaging is the only feasible damage assessment possibility!



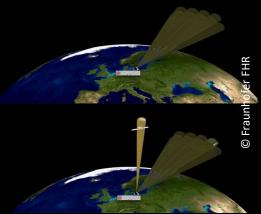


Detection and Cataloguing of Space Debris

GESTRA – German Space Surveillance Radar

- Federal government outlines necessity of establishing a national competence center for documentation and evaluation of the current situation in space
- Quasi-monostatic pulsed phased array radar in L-band (~ 1.3 GHz)
- Surveillance in orbital heights of 300-3000 km
- Electronic swiveling through HF-based modification of the wave front during transmission or digital beam forming during reception
- Numerous operating modes (surveillance volumes vs. detection capacity, track-while-scan, etc.)







Detection and Cataloguing of Space Debris

GESTRA – German Space Surveillance Radar

- GESTRA consists of:
 - 2 containers one receiving (RX) and one transmitting (TX) unit
 - both subsystems have independent infrastructure (energy, cooling, climate control, etc.)
 - size each: 18 m x 4 m x 4 m; weight about 90 t
- each unit (RX&TX) contains:
 - a phased array antenna with 256 individual elements
 - mounted on top of a 3-D positioner

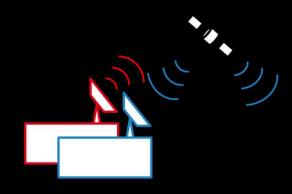




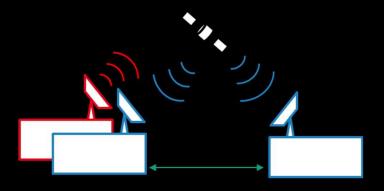
Further detection concepts and projects

GESTRA EU-SST & GESTRA networks

- the energy reflected away can be collected by a bi-static sensor configuration as implemented by EUSST (lower graph)
- alternatively, larger search volumes can be "scanned" simultaneously
- more accurate tracks can be calculated by combining multiple sensor contributions
- the sensitivity of the overall system is significantly increased by bi- or even multi-static arrangement of the sensors
- high-precision synchronization and appropriate signal processing are required for implementation



mono-static radar configuration



bi-static radar configuration



Further detection concepts and projects

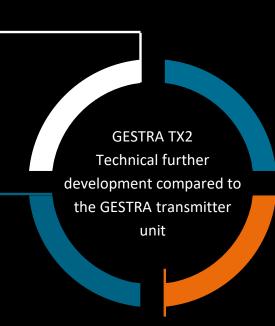
GESTRA TX2

Transmitting technology

- Development of a new transmitter module by means of the new semiconductor technology based on GaN
- Increase of the transmitting power by a factor of approx. 2

Networking

- Networked operation of TX2 together with GESTRA and EUSST (two transmitting and two receiving units)
- Time, phase and frequency synchronization
 - Performance improvement



Cooling

- Higher transmitting power requires new cooling development
 - Pulsating Heat Pipes-technology
- Project partners:Fraunhofer IPM & IWU

Signal generation

- New development of the Plank controller modules in the form of a multi-channel AWG
- Used for controlling the individual transmitter modules
 - Decentralized signal generation

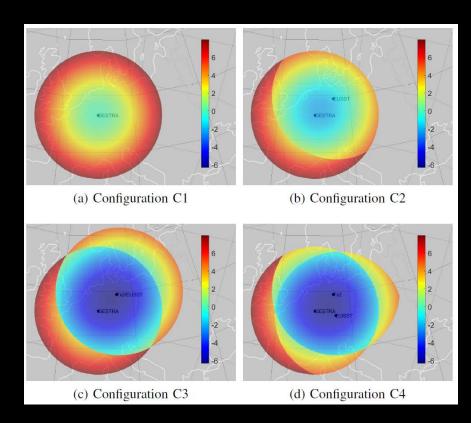


Further detection concepts and projects

GESTRA network configurations

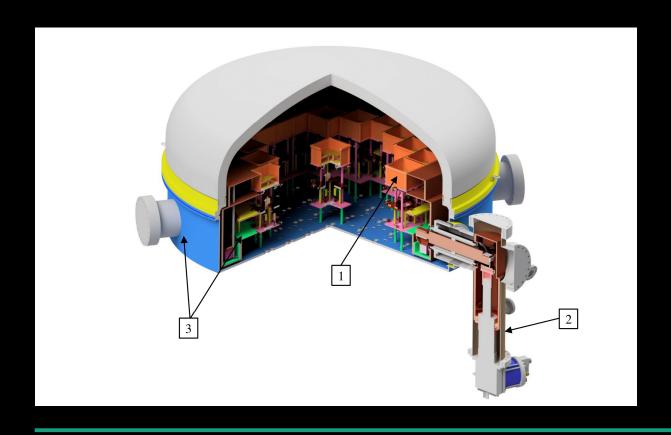
Improvement of the minimum detectable RCS in dB at 1,000 km altitude for different GESTRA network configurations.

- C1: GESTRA
- C2: GESTRA + GESTRA EUSST
- C3: GESTRA + GESTRA EUSST / GESTRA TX2
- C4: GESTRA + GESTRA EUSST + GESTRA TX2
- Attention: RCS values are normalized to the minimum detectable RCS of GESTRA (0 dB)
- Results:
 - C2: Detection power increases by up to 2.3 dB
 - C3/C4: Detection power increases by up to 6,3 dB
- Detection of targets that are up to four times smaller





Cryo-Cooled 37-Element Phased Array Radar Receiver



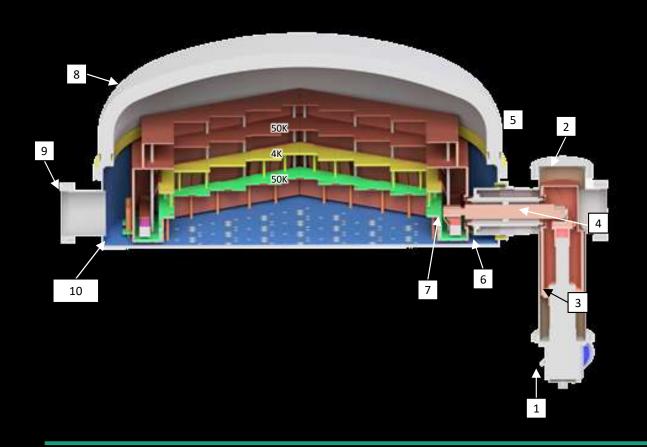
Components:

- 1 RF-Unit-Cell
- 2 2-Stage Cryocooler
- 3 Vacuum Vessel

Diameter 1.5 m



Cryo-Cooled 37-Element Phased Array Radar Receiver



Vaccum Vessel with Cryocooler

- 1 Sumitomo Cryocooler
- 2 Vacuum Vessel
- 3 50K-Stage
- 4 4K-Stage
- 5 Copper-Shield
- 6 Vacuum Barrier
- 7 Copper Rod
- 8 Radom
- 9 Vacuum Pump Interface
- 10 Lower Flange



Some simulation results

Maximum deformation along the Z axis was 10.45 mm

-> To reduce this value, either the thickness of the vacuum vessel can be increased or the 4 stand feet can be scaled higher

Temperature distribution in the LNA area was between 23.5 and 29.7 K and in the antenna area between 108.4 and 134.2 K

- -> Temperature differences can be reduced by increasing the cross-sectional area
- Based on the simulation results, optimizations can be made and incorporated into the hardware structure
- A. Froehlich *et al* 2022 *IOP Conf. Ser.: Mater. Sci. Eng.* **1240** 012102



Conclusion

- Showed Space Debris is a rising problem for our infrastructures and services
- TIRA allows to support all stages of space missions
- GESTRA family a novel phased array systems for detection and catalogue space debris
- New concepts of GESTRA systems such as GESTRA Networks or cryocooled receivers improve the performance further





Thank you for your attention

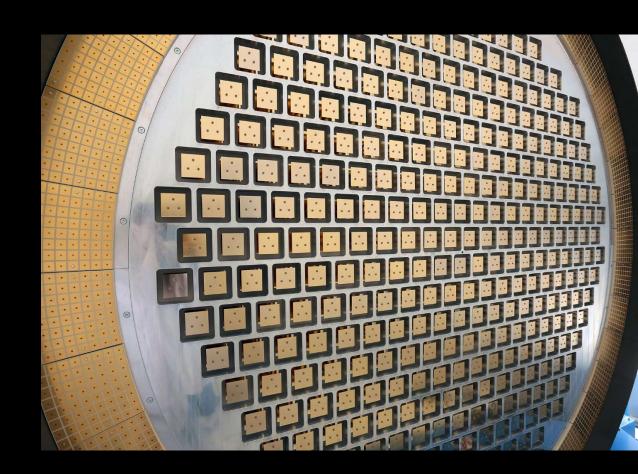
Thanks to all colleagues from Fraunofer

Contact:

Andreas Froehlich

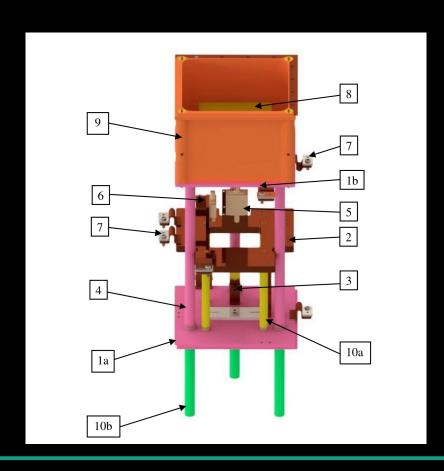
andreas.froehlich@fhr.Fraunhofer.de

Fraunhofer Institute for High Frequency
Physics and Radar Techniques FHR, Germany





Cryo-Cooled 37-Element Phased Array Radar Receiver



RF unit cell:

- 1 50K-Plate
- 2 4K-Plate
- 3 Cooper Rod for RF-Cables
- 4 Copper Spacer
- 5 LNA (Low noise amplifier)
- 6 LNA Fixing Plate (Copper)
- 7 Copper Stands
- 8 Patch-Antenna
- 9 Cavity
- 10 GFK-Spacer



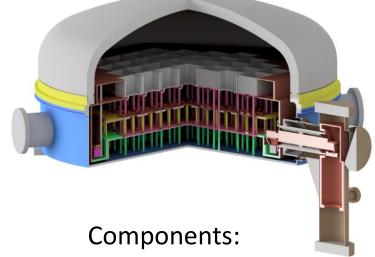
Preparing and Optimizing of 3D-Simulation Model





Interfaces: 6.648

Simplified Model



1.467

Interfaces:

3.382

Simulation Model and Project Scheme

■ Thermal Model

• Heat flow of the cables

Temperature in K	heat flow in W	
	Cable 1	Cable 2
4 50 300	8,14 10,42 17,82	5,89 5,86

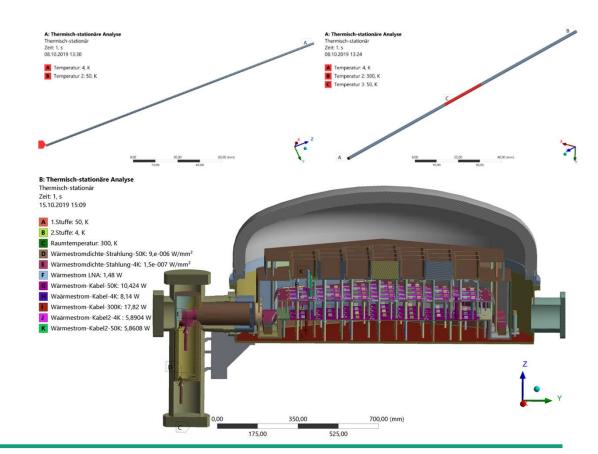
Heat flow of the LNAs

$$\dot{Q}_{Ges} = n \cdot k \cdot \dot{Q}_{LNA} = 1,48 \text{ W}$$

• Heat transfer through thermal radiation

$$\dot{q}_{300-50\text{K}} = 9\text{E} - 006 \text{ W/mm}^2$$

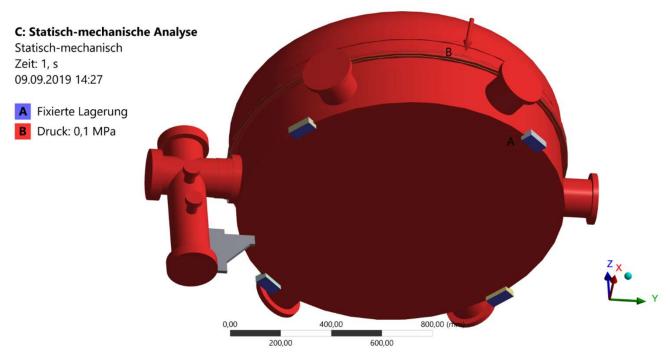
 $\dot{q}_{50-} = 1.5\text{E} - 007 \text{ W/mm}^2$





Simulation Model and Project Scheme

Mechanical Model



- Fixed storage on the 4 feet of the vacuum vessel
- Atmospheric pressure in external surfaces

Simulation Model and Project Scheme

Modal Analysis

■ Determination of the eigenfrequencies by the eigenvalue equation

$$(-\omega_i^2 \quad [M] + [K])$$

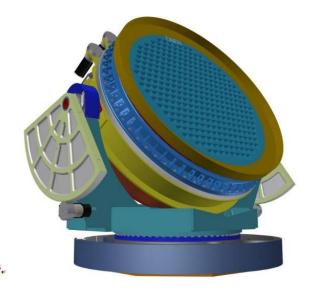
Harmonic Analysis

- Response of the system at the determined natural frequencies
- 1 Newton stimulating force

Response Spectrum

- Determination of the deformation for a given response spectrum
- At a maximum angular velocity of 5 °/s
- A maximum Angular acceleration of 5 ° / s2 results in a total acceleration of

$$a_{ges} = \sqrt{a_r^2 + a_t^2} = 61,31 \frac{mm}{s^2}$$

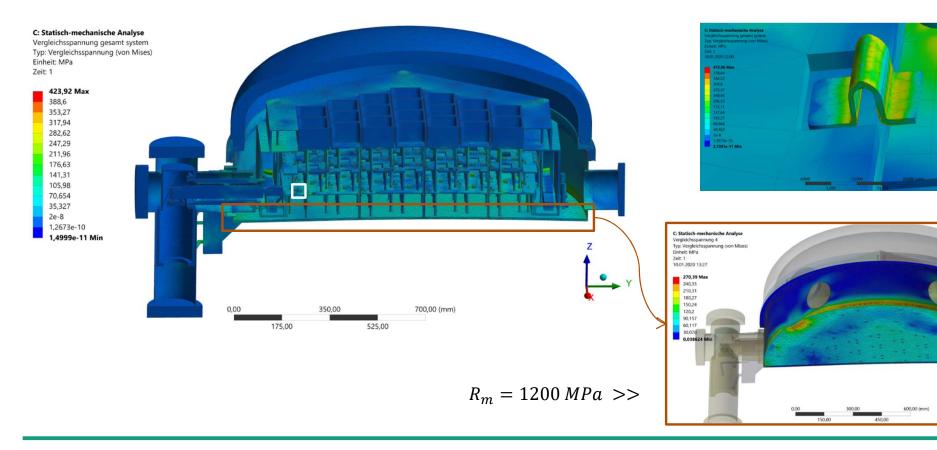


- PSD-Analysis (only Paper)
 - Determination of the deformation 3σ (99,7%)

PSD =
$$\frac{a^2_{ges}}{\Delta f}$$
 = 21,57 mm²s⁻⁴/Hz

Result and Evaluation

Mechanics Model





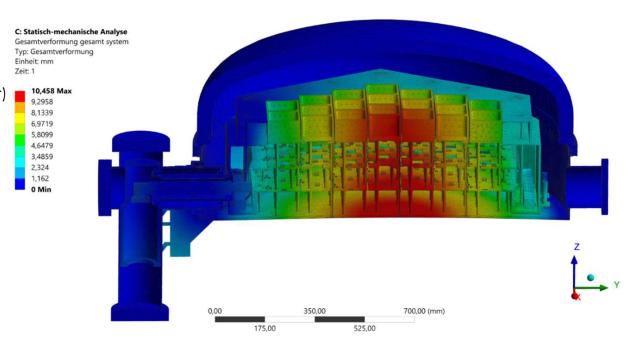
Mechanical Model

Deformation

- Maximal deformation of the vacuum vessel because of pressure difference (upper) and shrinking factor at 4 Kelvin
- Verification by analytical calculation

$$f = \frac{0,171 \cdot p \cdot r^4}{E \cdot h^3} = 15,2 \ mm$$

 Deviation due to feet fixation on the vacuum vessel

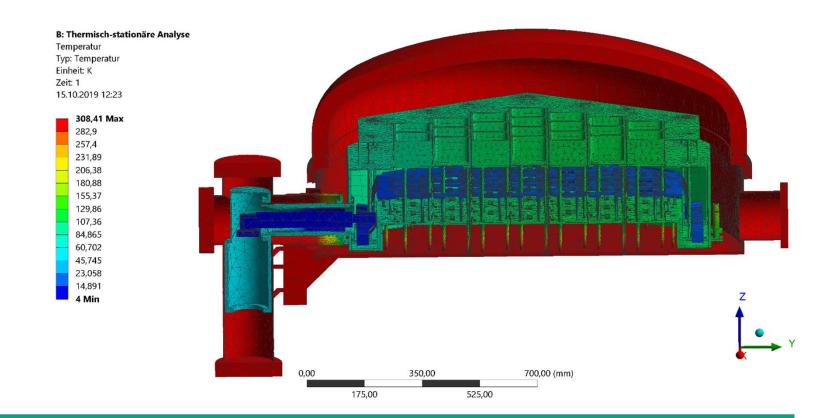


Result and Evaluation

Thermal Simulation

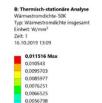
- LNA area: 23-29 K
- Antenna area:108-134 K
- Optimization:

$$\Delta T = \frac{d \cdot \dot{Q}}{A \cdot \lambda}$$



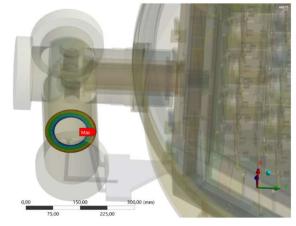


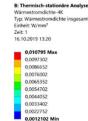
Thermal Simulation

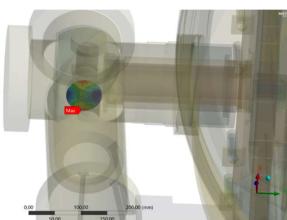


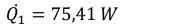
0,0047072

0,0037346









$$\dot{Q}_2 = 10,36 W$$

RDK-415D Typical Load Map (50Hz)



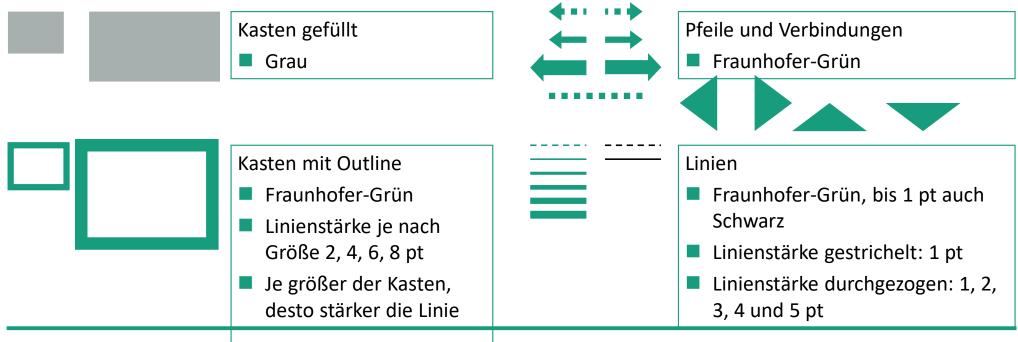
$$T_1 = 105 K$$
 $T_2 = 9 K$

$$T_2 = 9 \, K$$

Grafische Elemente

Kästen, Pfeile, Verbindungen und Linien (Auswahl) ! DIESE FOLIE AUS FINALER PRÄSENTATION LÖSCHEN!

folgende Elemente können hier per Rechtsklick kopiert und an gewünschter Stelle in der neuen Präsentation per Rechtsklick wieder eingesetzt werden:



Farben

! DIESE FOLIE AUS FINALER PRÄSENTATION LÖSCHEN !

- folgende Farben können über die Powerpoint-Farbauswahl hier aufgenommen und damit in der neuen Präsentation angewendet werden:
- Überschriften / Fließtext / Quellenangaben / Bildunterschriften / Grafikauszeichnungen
- Grafikauszeichnungen
- Aufzählungen / Nummerierungen erster Ebene / grafische Elemente
- Grafiken
- Fonds hinter Grafiken

