

ALPACA: The Advanced L-band Phased Array Camera For Astronomy

Mitchell C. Burnett, Karl F. Warnick, and Brian D. Jeffs Department of Electrical and Computer Engineering Brigham Young University, Provo, UT, USA

Amit Vishwas, Stephen Parshley, George Gull, and Donald Campbell Cornell Center for Astrophysics and Planetary Science Cornell University, Ithaca, NY, USA Phased Array Feeds & Advanced Receivers Workshop 2022

BYU Electrical & Computer Engineering IRA A. FULTON COLLEGE OF ENGINEERING

ALPACA Project **Overview**

ALPACA Project

- National Science Foundation Award AST-1636645
- Brigham Young University and Cornell University
	- 4 years, \$5.8M, started July 2018
	- Currently in 5th year, commissioning expected July 2023
- Arecibo Telescope
	- Decommissioned and then Collapse on Dec. 1, 2020
	- Evaluation and funding efforts under way for transition to the Green Bank Telescope
- Cornell: PAF front-end, including electronics, array elements, dewar, cryogenics, mechanical engineering
	- Subaward to ASU for LNA
- Brigham Young University and Cornell University
• 4 years, \$5.8M, started July 2018
• Currently in 5th year, commissioning expected July 2023
• Arecibo Telescope
• Decommissioned and then Collapse on Dec. 1, 2020
• Eva F-engine firmware, digital beamformer, and data handling

ALPACA Overview

ALPACA Science Cases

RADIO ASTRONOMY

ALPACA Cryostat

RFoF

RF-over-Fiber Link

RADIO ASTRONOMY

RFoF, Cryostat Interface

RFoF, Link Performance

w/ the 4-ch TX r3.6, ch B, and the 16-ch RX rev2, ch 2

Digital Back End

Digital Back End, Architecture

Digital Back End, Specifications (1)

Digital Back End, Specifications (2)

Digital Back End, F-engine

- Xilinx ZCU216 board
- 16 on-chip 14-bit 2.5 Gsamp/sec ADCs, w/ digital down conversion, decimation, and lowpass filtering to complex baseband.
- On-chip ARM A53 processor
- 4x25 SFP28 network ports support up to a 100 Gigabit I/O data rate to HPC/GPUs.
- 12 of these boards w/ 12 inputs each to be used to support the 138 antenna
- Will directly sample RF over fiber downlinks with no analog mixer.

F-engine, Oversampled PFB

- comparing a fine "zoom" frequency response of 2048 $\frac{g}{g}$
point critically sampled and 4/3 oversampled PFB
(both use 8-tap branch filters)
When using the oversampled PFB: point critically sampled and 4/3 oversampled PFB $\frac{20}{\pi}$ -20 (both use 8-tap branch filters) **F-engine

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Separation**

Separation of 2048

F-engine
 $\frac{1}{2}$

omparing a fine "zoom" frequency response of 2048

Footh use 8-tap branch filters)

Then using the oversampled PFB:

Then using the overs **F-english**

F-engl

F-engl

omparing a fine "zoom" frequency response of 2

oint critically sampled and 4/3 oversampled PFE

both use 8-tap branch filters)

Vhen using the oversampled PFB:

• <u>Flat</u> in-channel passband re
- When using the oversampled PFB:
	-
	-
- Resulting in:
- when the spectral leakage to adjacent channels

No aliasing near coarse channel band edges
	- No aliasing near coarse channel band edges

F-engine, Oversampled PFB

- Hardware outputs from an ALPACA specified OSPFB (2048 branches, 8 polyphase taps per branch)
- Followed by a software second stage critically sampled PFB (32 point, 8 polyphase taps).

Digital Back End, XB-engine

25x 2U Tyan Transport HX, AMD EPYC 96 GB DDR4 RAM, PCIe 4.0

50x (2 per server) NVIDIA A10 GPU and ConnectX-5 dual-port InfiniBand EDR NIC

XB-engine Software Block Diagram

- Hashpipe processing pipeline
- Three primary tasks:
	- Capture network packets
		- ibverbs
	- Implement data processing modes
		- Calibration Correlator
		- Beamformer (Coarse frequency Transient mode)
		- Beamformer + Second stage channelizer (Fine frequency HI/spectral line mode)
	- Format and send output
- Communication through shared memory KV buffers
- Coordination among nodes managed by back end portion of M&C

Antenna Y Factor and Radiation Efficiency Measurement

Measuring Radiation Efficiency

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- Measurement Methods
	- Wheeler Cap
	- Anechoic Chamber
	- Radiometric (extended hot/cold RF sources)
		- Two antenna, with reference $\eta_{\text{rad}} = 1$ T_{ext}
		- Link Budget T_{rec}
		- Heated T_p
		- Use the antenna Y factor and standard Y factor measurement of Radiometer to obtain T_{rec} (this experiment)

$$
T_{\rm sys} = \eta_{\rm rad} T_{\rm ext} + \underbrace{(1 - \eta_{\rm rad}) T_{\rm p}}_{T_{\rm loss}} + T_{\rm rec}
$$

Measuring Radiation Efficiency

• The antenna Y factor is the ratio of noise powers at the system output when exposed to extended hot/cold loads

$$
Y = \frac{P_{\text{hot}}}{P_{\text{cold}}} = \frac{\eta_{\text{rad}} T_{\text{hot}} + T_{\text{loss}} + T_{\text{rec}}}{\eta_{\text{rad}} T_{\text{cold}} + T_{\text{loss}} + T_{\text{rec}}}
$$

• From the antenna Y factor, an equivalent active antenna temperature is computed as (Kerr 1999)

$$
T_{\text{eq}} = \frac{T_{\text{hot}} - Y T_{\text{cold}}}{Y - 1}
$$

• If we move the reference plane for the system noise temp to before antenna losses (e.g., to an equivalent sky temperature)

$$
T_{\rm sys} = \eta_{\rm rad} T_{\rm ext} + \underbrace{(1 - \eta_{\rm rad}) T_{\rm p}}_{T_{\rm loss}} + T_{\rm rec}
$$
\n
$$
\frac{T_{\rm sys}}{\eta_{\rm rad}} = T_{\rm ext} + \underbrace{T_{\rm loss} + T_{\rm rec}}_{T_{\rm rad}}
$$
\n
$$
\eta_{\rm rad} = \frac{(T_{\rm p} + T_{\rm rec})(Y - 1)}{T_{\rm hot} - T_{\rm p} + Y(T_{\rm p} - T_{\rm cold})}
$$

Measuring Radiation Efficiency

Ground shield with absorber foam (hot load)

Extrapolation to Cryogenic Temperatures

- At cryogenic temperatures, electrical conductivity is quantified by relative resistivity ratio (RRR)
- ALPACA dipole is fabricated primarily from Aluminum 6061
- coax core that runs from the wedge arms to SMA
- micro-in (2.54 um) thickness and a layer of copper for adhesion
- At L band, the gold plating is several skin depths thick, so the gold substantially determines the electrical conductivity of the antennas.
- Using a value of 12 for the RRR (Finger & Kerr 2008) of gold at 20 K we extrapolate the antenna radiation efficiency and estimated loss contributed by the antenna

$$
\eta_{\rm rad, 20 \, K} = \frac{R_{\rm rad}}{R_{\rm rad} + R_{\rm loss}/{\rm RRR}} \qquad T_{\rm loss} = (1 - \eta_{\rm rad})T_{\rm p}
$$

CASPER Integration

CASPER Integration

- Gen1 and Gen 3 RFSoC Platforms:
	- RFSoC $4x2$, ZCU216, ZCU208 RFSoC
	- **ZCU111, ZRF16**
- RFSoC ADC and DAC
- Generalized OSPFB
- URAM Packetizer
- 100GbE Core

CASPER Integration

RFSoC Tutorials: https://caspertoolflow.readthedocs.io/projects/tutorials/en/latest/

- From Getting started with RFSoC
- To using the RFDC
- And how to stream ADC samples using 100 GbE and capture data packets at a NIC for processing

Thank you!

RADIO ASTRONOMY

CASPER Integration

- 4 element ULA
- System Generator Polyphase DDS for generating 2 signals from different directions
- Complex multiply blocks for weight application
- Beamforming weights computed in python
- PFB-based spectrometer

- SOI tone sweep $\frac{applied by the development of the image.}$ MHz at 20° and μ
- Interferer tone random FSK
from $260-410$ MHz at 70°

XB-engine

VLA RFI monitoring data from: http://www.vla.nrao.edu/cgi-bin/rfi.cgi

Using W8Mon for Mar 24 5-6 PM MST

SIDEREAL TIME FOR GIVEN DATE, TIME AND LONGITUDE

Local Date = 20220324 = 2022 Mar 24 [Thursday] Gregorian 1000 $= 2459662.5 = JD number for $00h$ TT of date$

Local Time = $17:23:00 = 0.724305555556$ day T Zone Diff = $+06:00 = +0.250000000000$ day Delta T = $+00:00:00 = +0.0000000000000$ day

Full astronomical JD number = Sum of all given date/time elements above $JD = 2459663.474305555556 = (JD00 + LTFrac + TZFrac + dTFrac)$

Date/Time TT Corresponding to Full JD 20220324 23:23:00.0 TT

Longitude = $-112^{\circ} 21' 07.2'' = -112.3520000000''$ (W) $= -07h$ 29m 24.5s = $-7.4901333333h$

Local Mean Sidereal Time at Longitude -112.35200000000 (W) 04h 03m 15s = 4.0540970100 h 60° 48' 41" = 60.8114551506°

Local True Sidereal Time at Longitude = -112.3520000000° (W) 04h 03m 14s = 4.0538593534 h 60° 48' 28" = 60.8078903011°

SIDEREAL TIME FOR GIVEN DATE, TIME AND LONGITUDE

Local Date = 20220324 = 2022 Mar 24 [Thursday] Gregorian $= 2459662.5 = JD$ number for 00h TT of date JD00

Local Time = $21:04:00 = 0.877777777778$ day
T Zone Diff = +06:00 = +0.2500000000000 day Delta T = $+00:00:00 = +0.000000000000$ day

Full astronomical JD number = Sum of all given date/time elements above $JD = 2459663.627777777778 = (JD00 + LTFrac + TZFrac + dTFrac)$

Date/Time TT Corresponding to Full JD 20220325 03:04:00.0 TT

Longitude = $-112^{\circ} 21' 07.2'' = -112.3520138889^{\circ}$ (W) $= -07h$ 29m 24.5s = $-7.4901342593 h$

Local Mean Sidereal Time at Longitude -112.3520138889° (W) 07h 44m 51s = 7.7475140431 h $116° 12' 46" = 116.2127106460°$

Local True Sidereal Time at Longitude = -112.3520138889° (W) 07h 44m 50s = 7.7472766704 h $116° 12' 33" = 116.2091500566°$

$T_{\rm sky} = T_{\rm Cont} + T_{\rm Cas\ A} + T_{\rm HI} + T_{\rm CMB}$

Assumptions from this…

- the shield, and contributions from the sun
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